
Light, Camera and Shading

CS 543 / ECE 549 – Saurabh Gupta

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<http://saurabhg.web.illinois.edu/teaching/ece549/sp2020/>

Overview

- Cameras with lenses
 - Depth of field
 - Field of view
 - Lens aberrations
- Brightness of a pixel
 - Small taste of radiometry
 - In-camera transformation of light
 - Reflectance properties of surfaces
 - Lambertian reflection model
 - Shape from shading

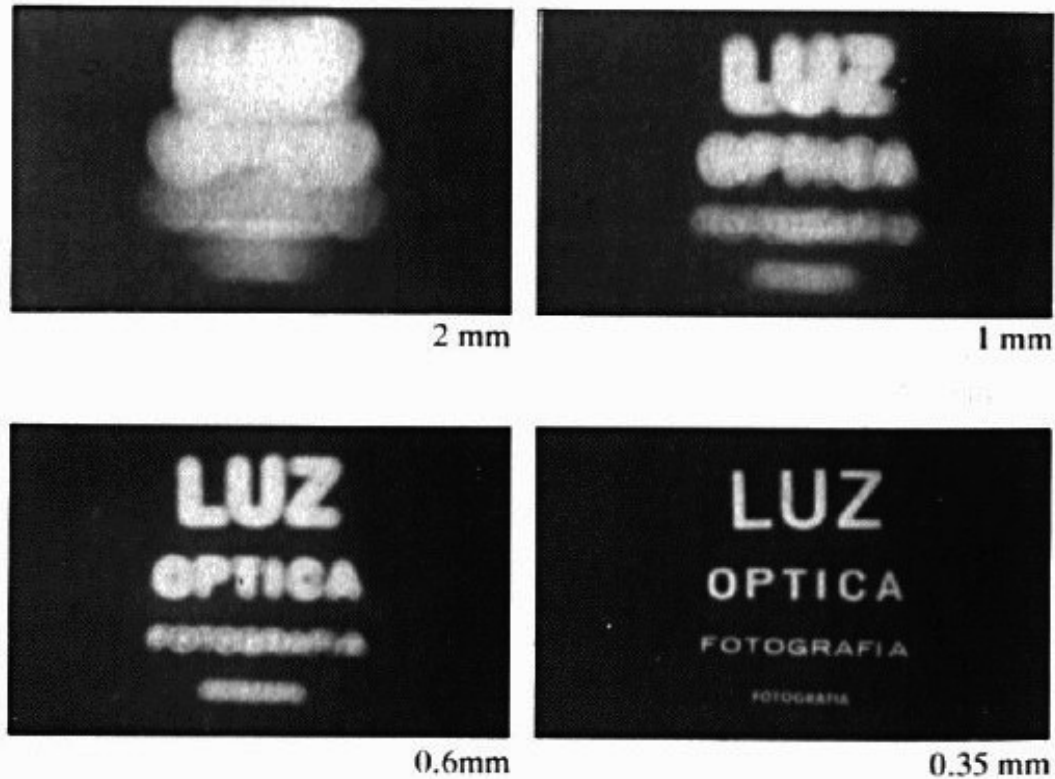
Building a Real Camera



Home-made pinhole camera



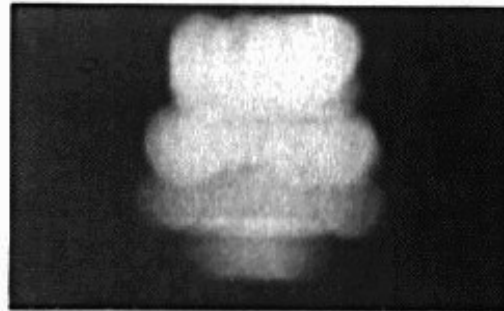
Shrinking the aperture



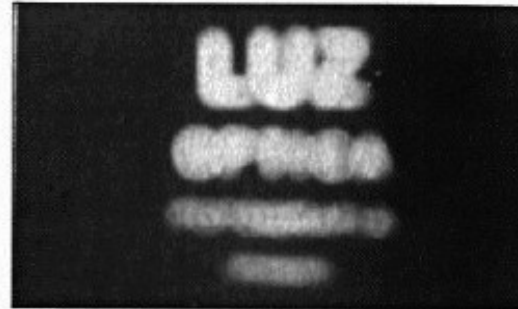
Why not make the aperture as small as possible?

- Less light gets through
- Diffraction effects...

Shrinking the aperture



2 mm



1 mm



0.6mm



0.35 mm



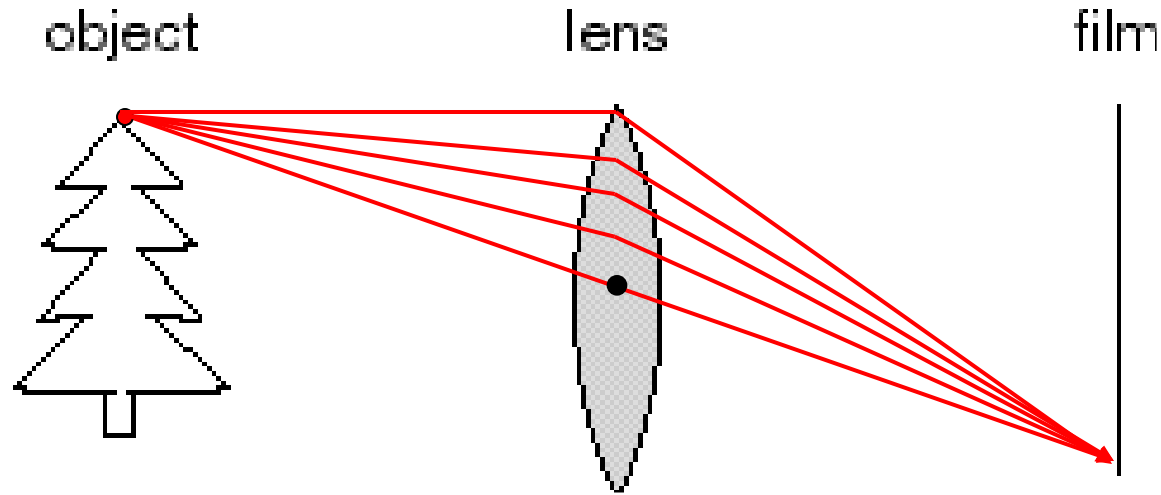
0.15 mm



0.07 mm

Adding a lens

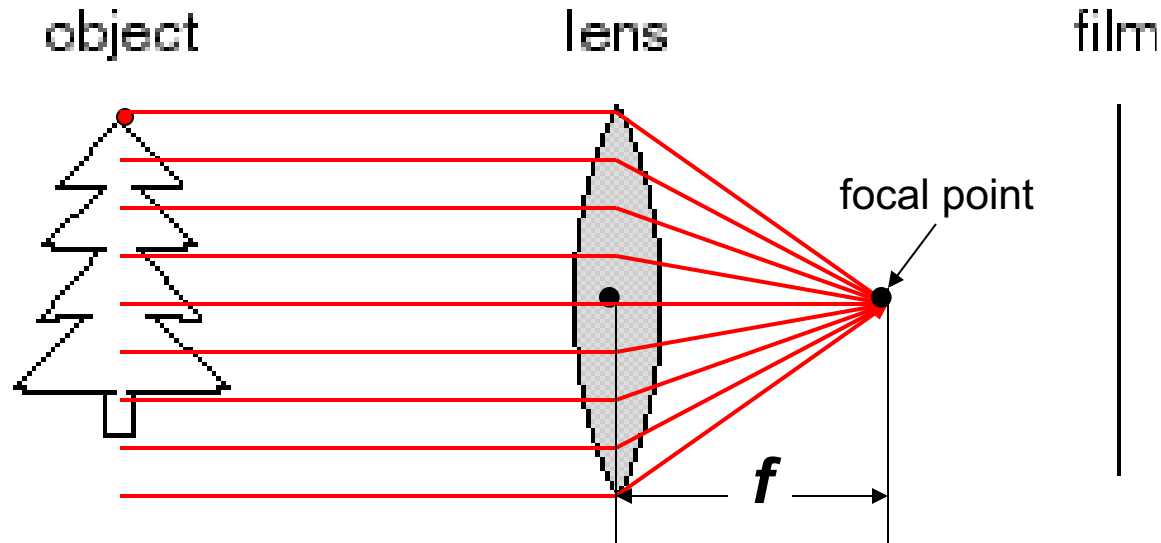
Adding a lens



A lens focuses light onto the film

- Thin lens model:
 - Rays passing through the center are not deviated (pinhole projection model still holds)

Adding a lens

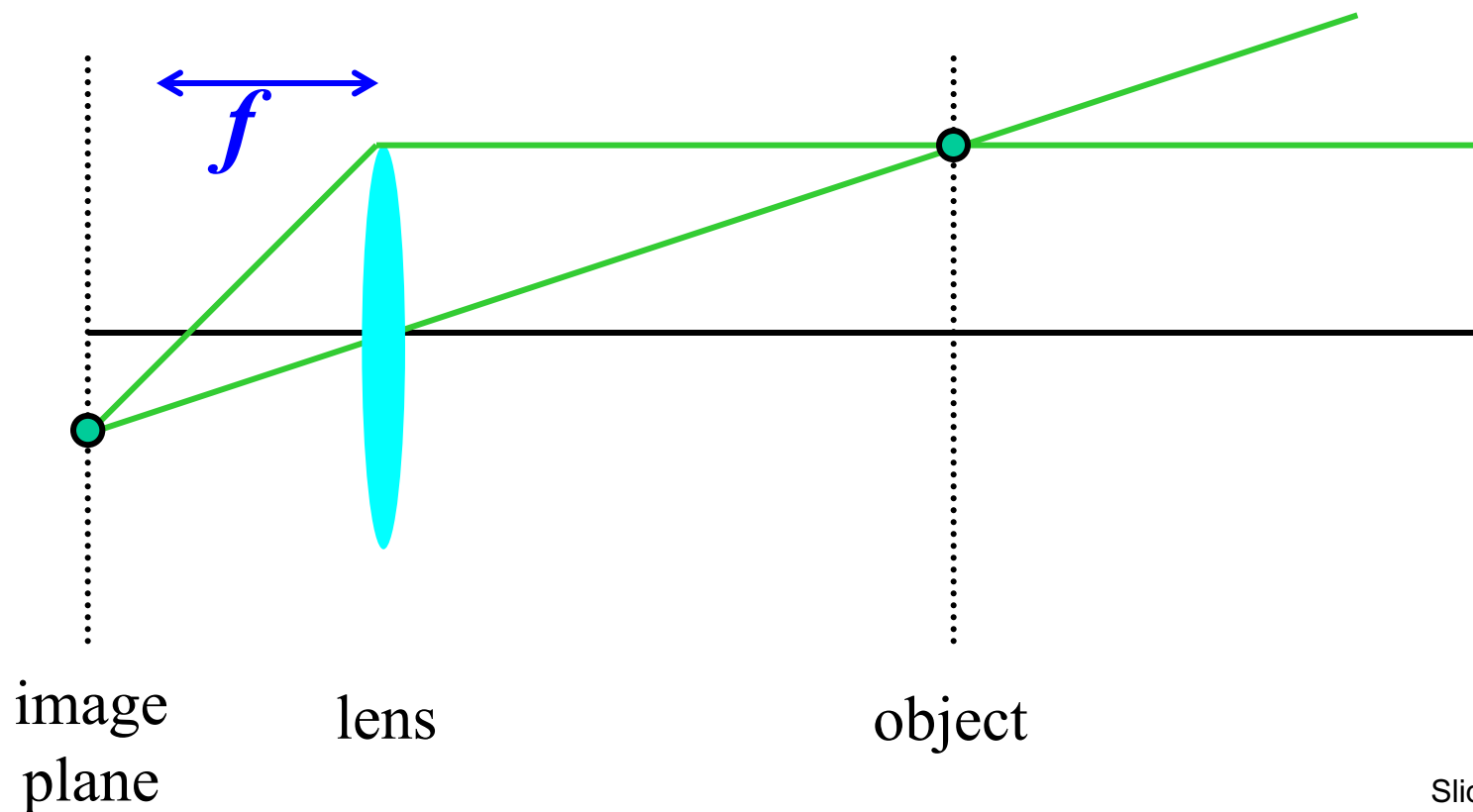


A lens focuses light onto the film

- Thin lens model:
 - Rays passing through the center are not deviated (pinhole projection model still holds)
 - All rays parallel to the optical axis pass through the *focal point*
 - All parallel rays converge to points on the *focal plane*

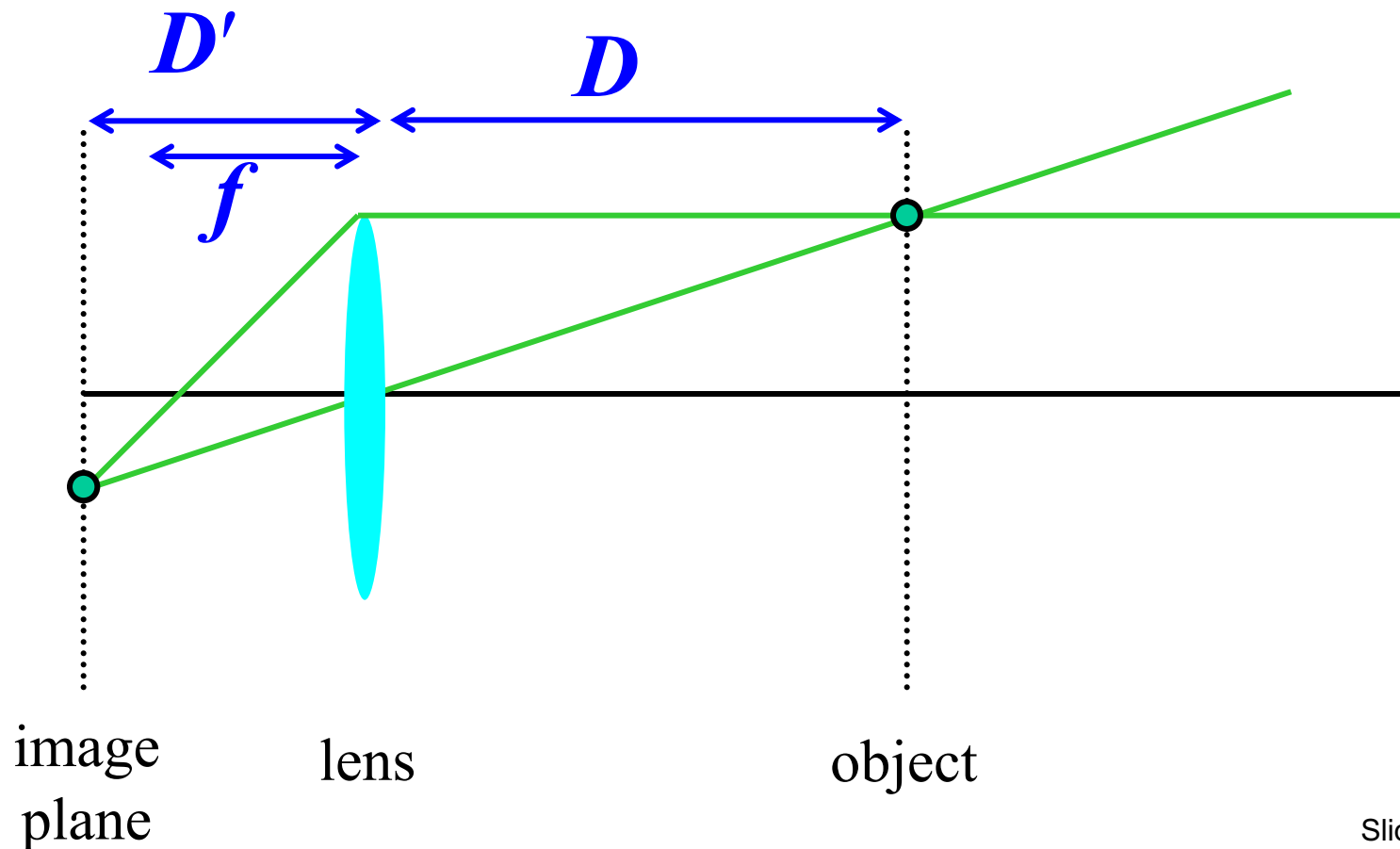
Thin lens formula

- Where does the lens focus the rays coming from a given point in the scene?



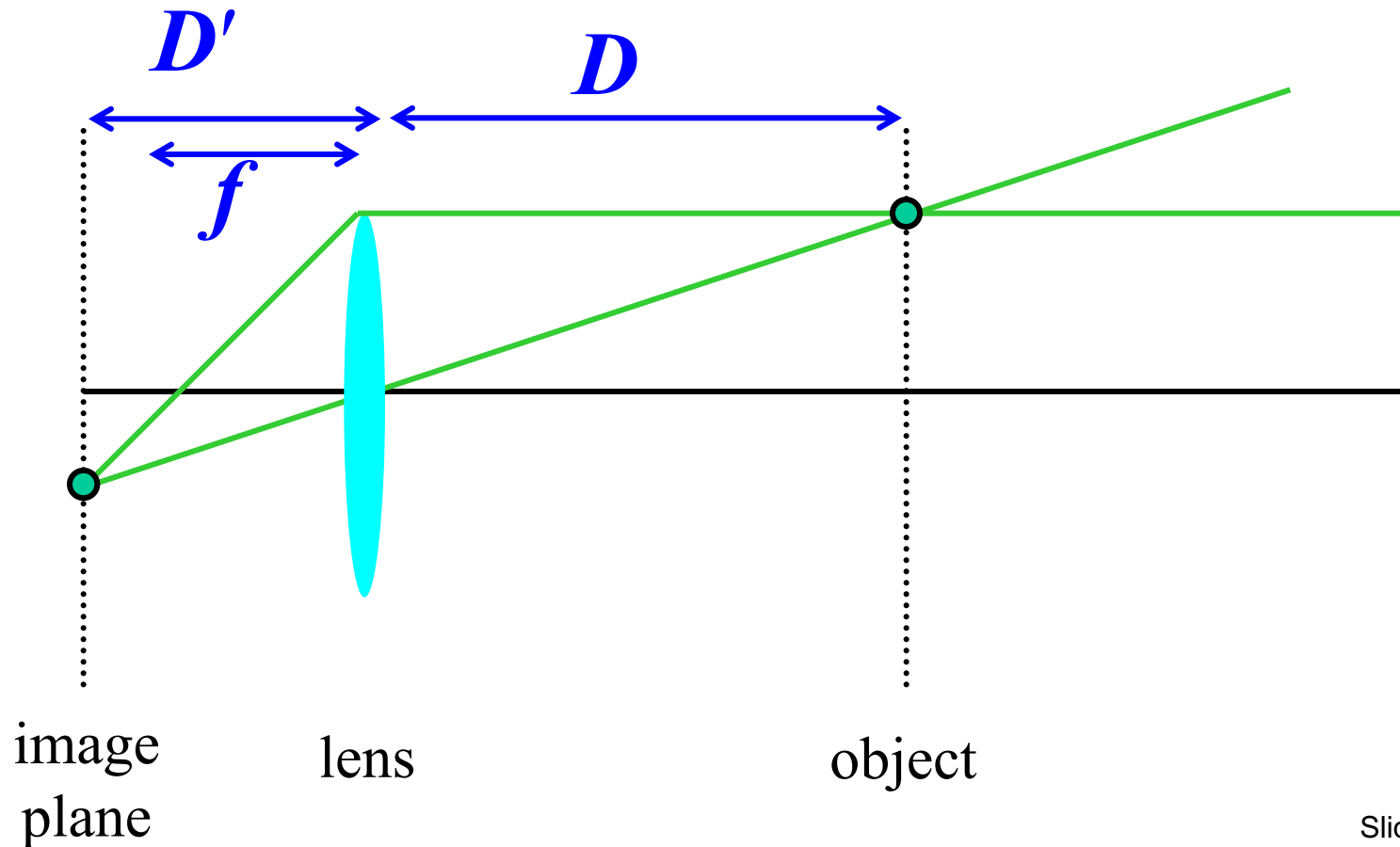
Thin lens formula

- What is the relation between the focal length (f), the distance of the object from the optical center (D), and the distance at which the object will be in focus (D')?



Thin lens formula

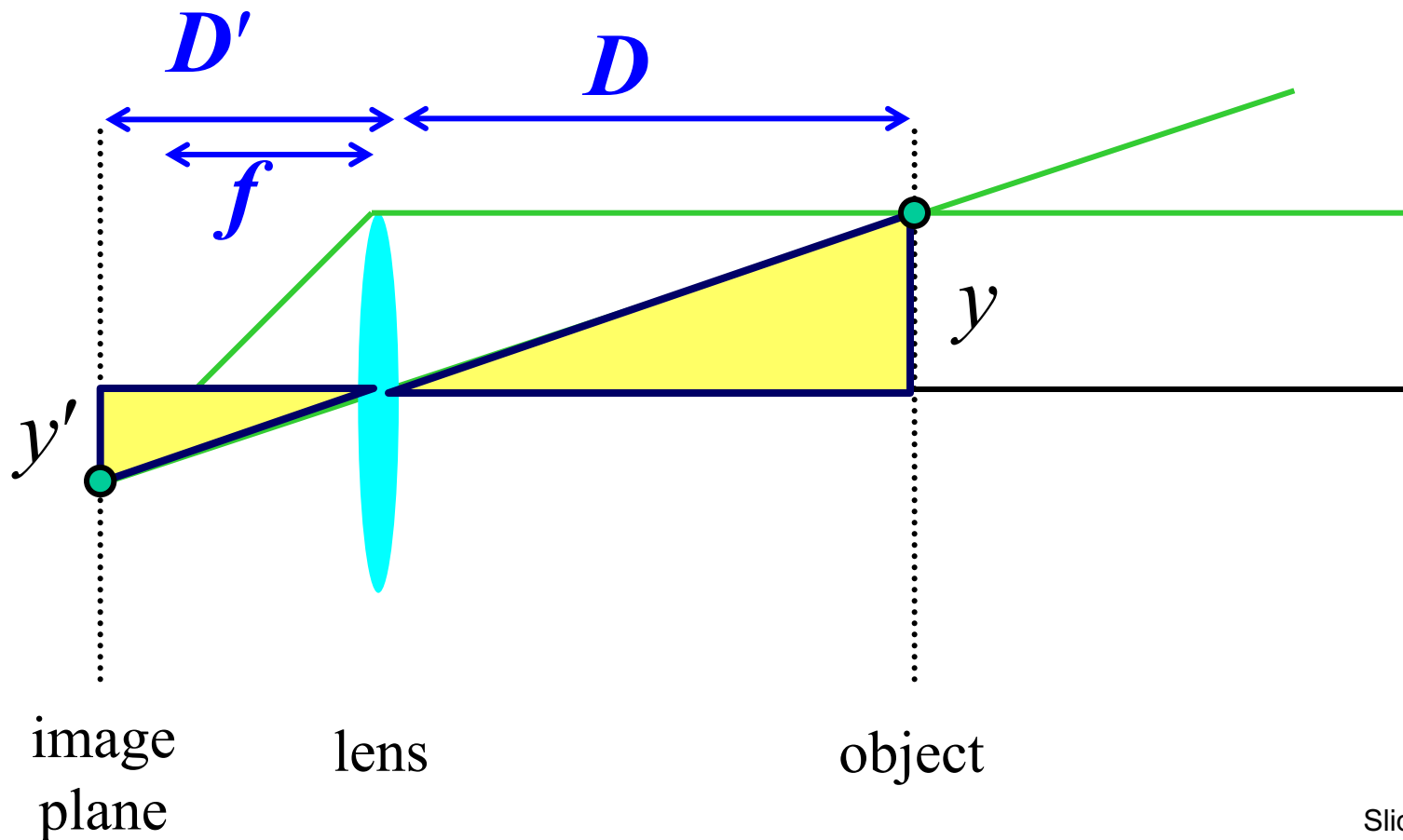
Similar triangles everywhere!



Thin lens formula

Similar triangles everywhere!

$$y'/y = D'/D$$

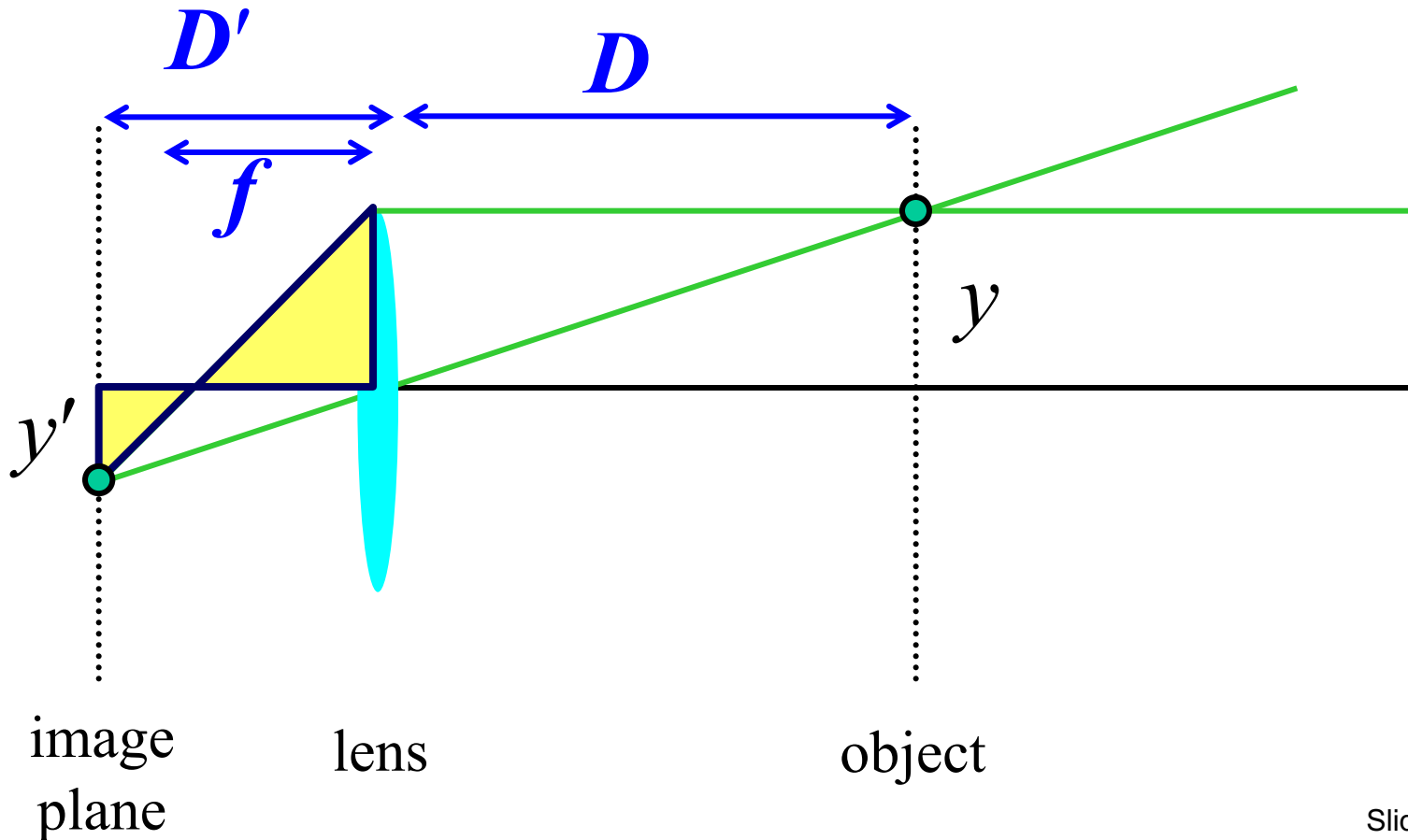


Thin lens formula

Similar triangles everywhere!

$$y'/y = D'/D$$

$$y'/y = (D' - f)/f$$

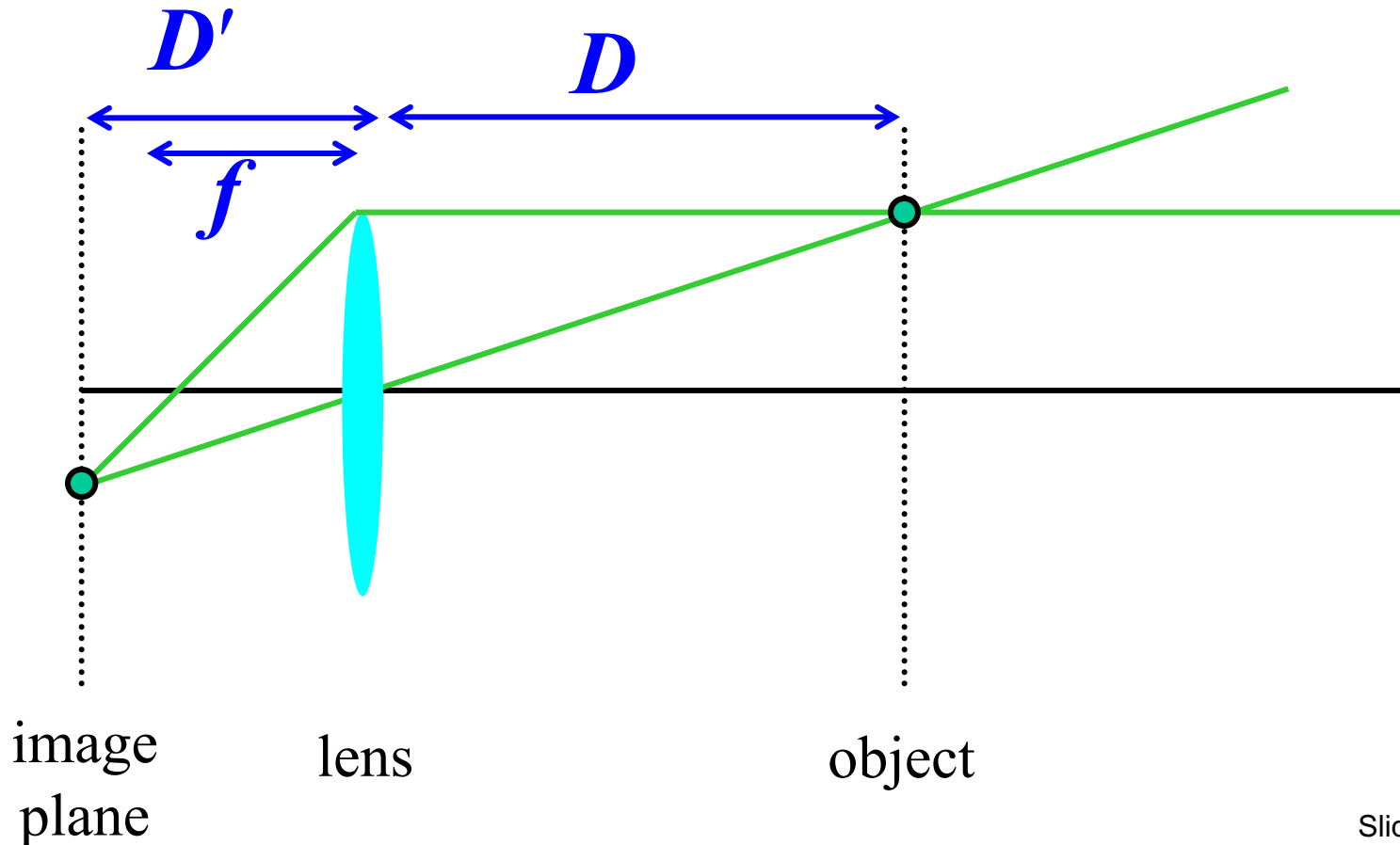


Thin lens formula

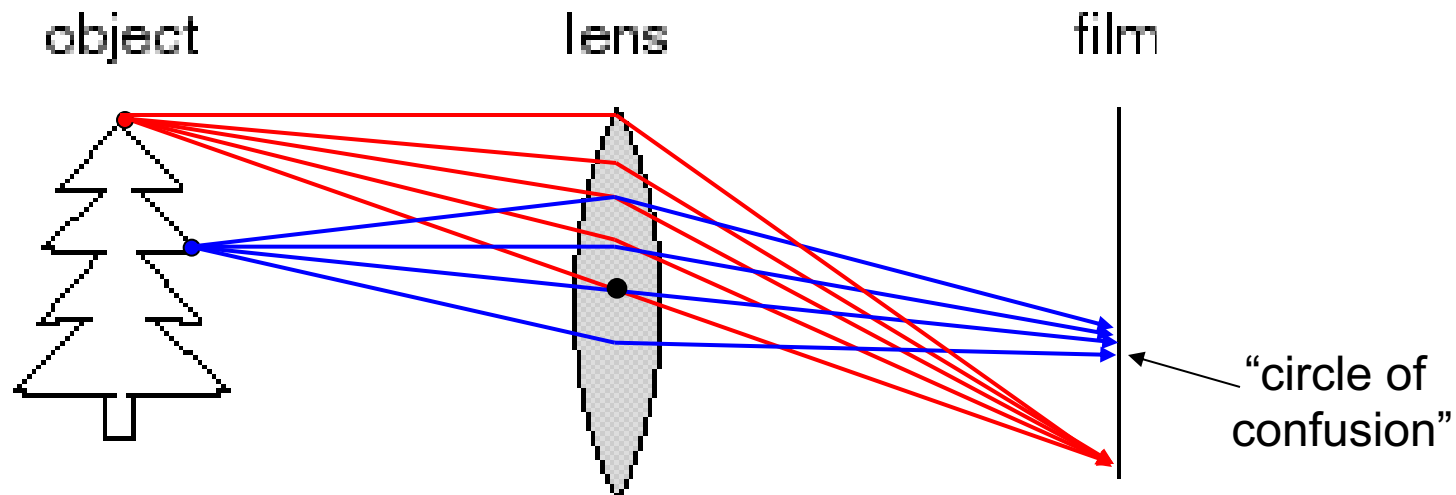
$$\frac{1}{D'} + \frac{1}{D} = \frac{1}{f}$$

Any point satisfying the thin lens equation is in focus.

What happens when D is very large?



Depth of Field



For a fixed focal length, there is a specific distance at which objects are “in focus”

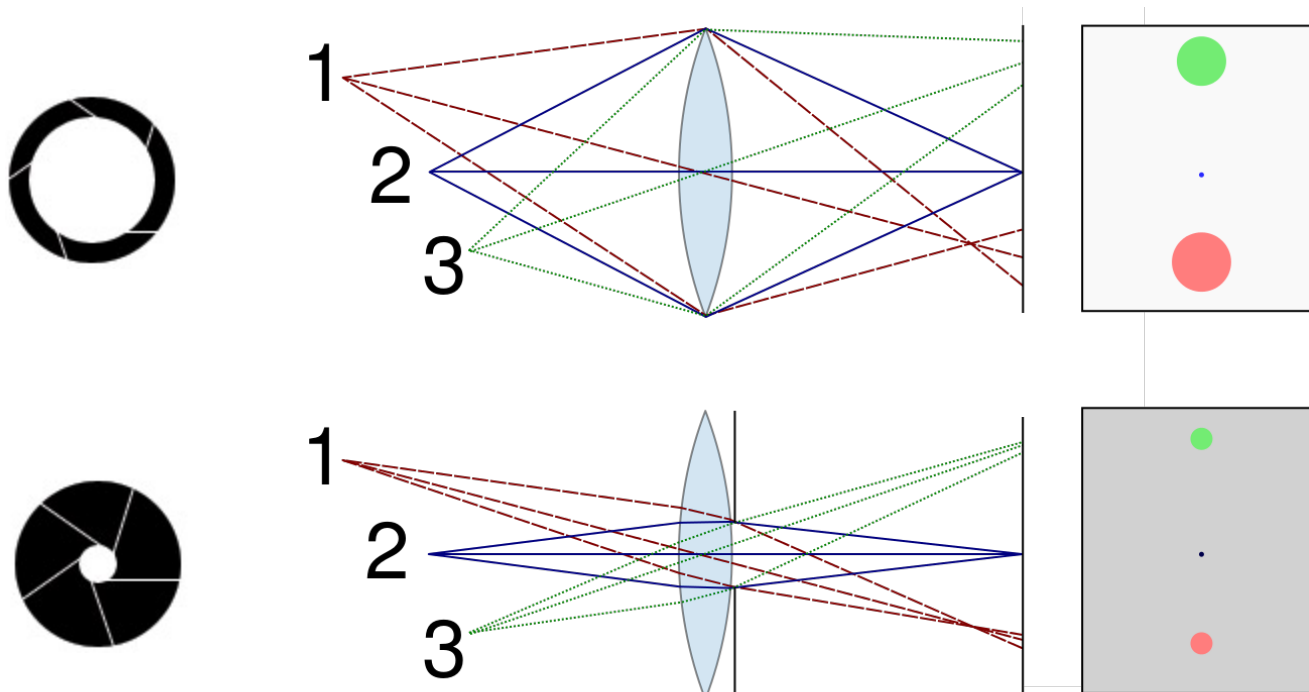
- Other points project to a “circle of confusion” in the image

Depth of Field



<http://www.cambridgeincolour.com/tutorials/depth-of-field.htm>

Controlling depth of field



Changing the aperture size affects depth of field

- A smaller *aperture* increases the range in which the object is approximately in focus
- But small aperture reduces amount of light – need to increase *exposure*

Varying the aperture

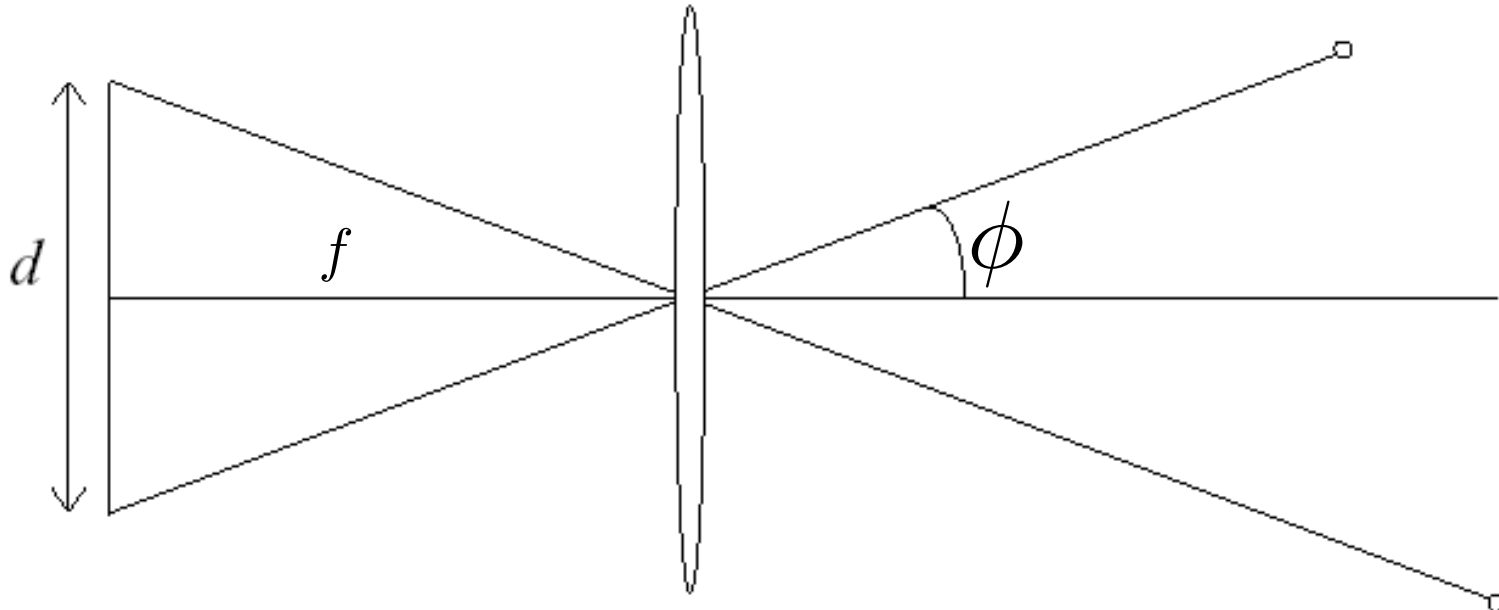


Large aperture = small DOF



Small aperture = large DOF

Field of View

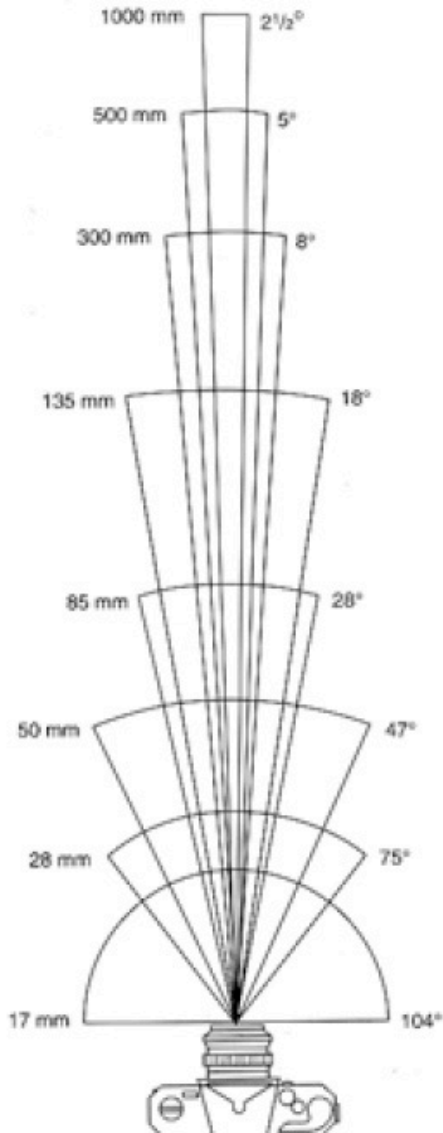


FOV depends on focal length and size of the camera retina

$$\phi = \tan^{-1}\left(\frac{d/2}{f}\right)$$

Larger focal length = smaller FOV

Field of View



17mm



28mm

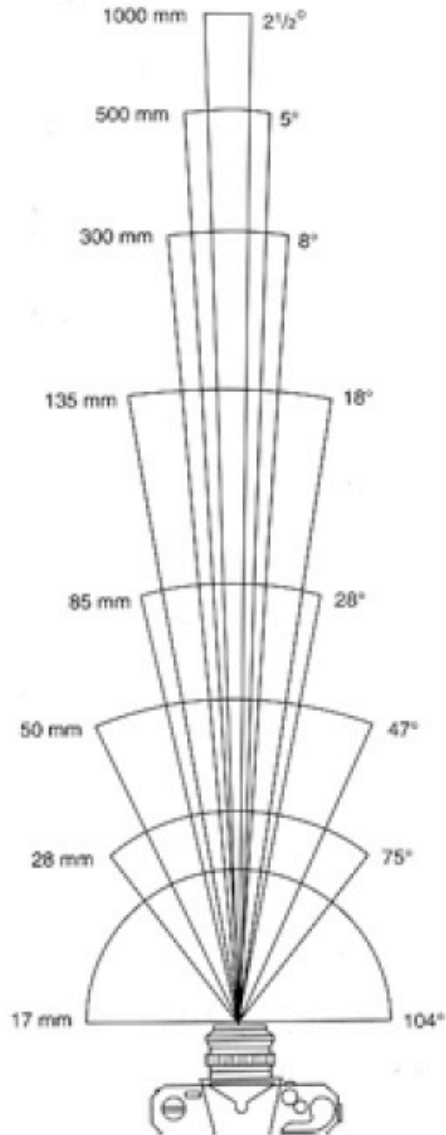


50mm

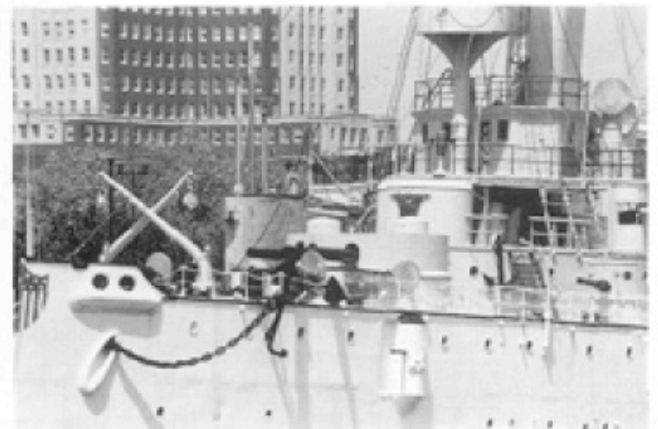


85mm

Field of View



135mm



300mm

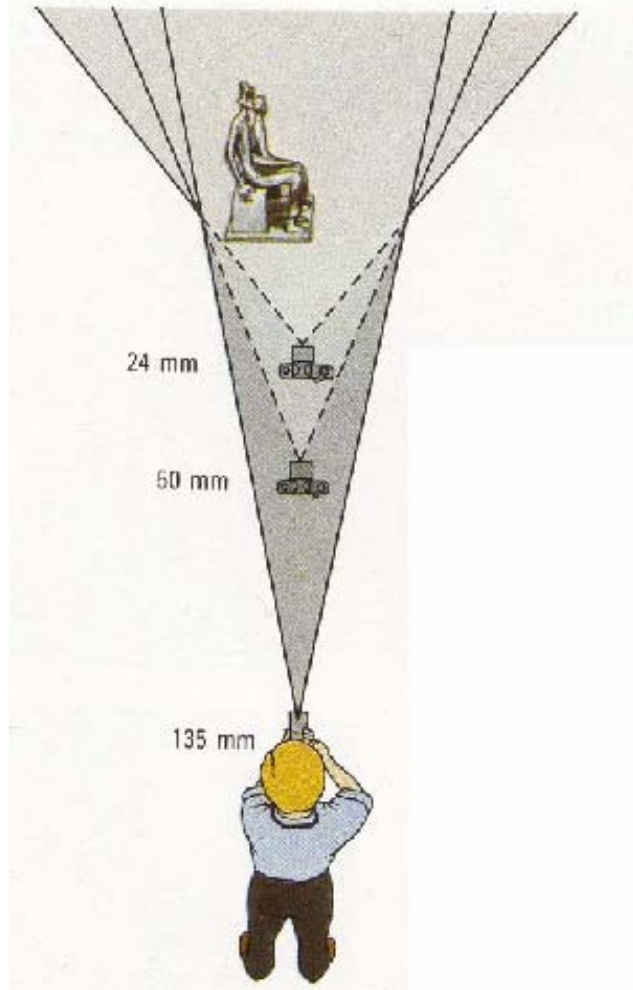


500mm



1000mm

Field of View / Focal Length



Large FOV, small f
Camera close to car



Small FOV, large f
Camera far from the car

Same effect for faces



wide-angle

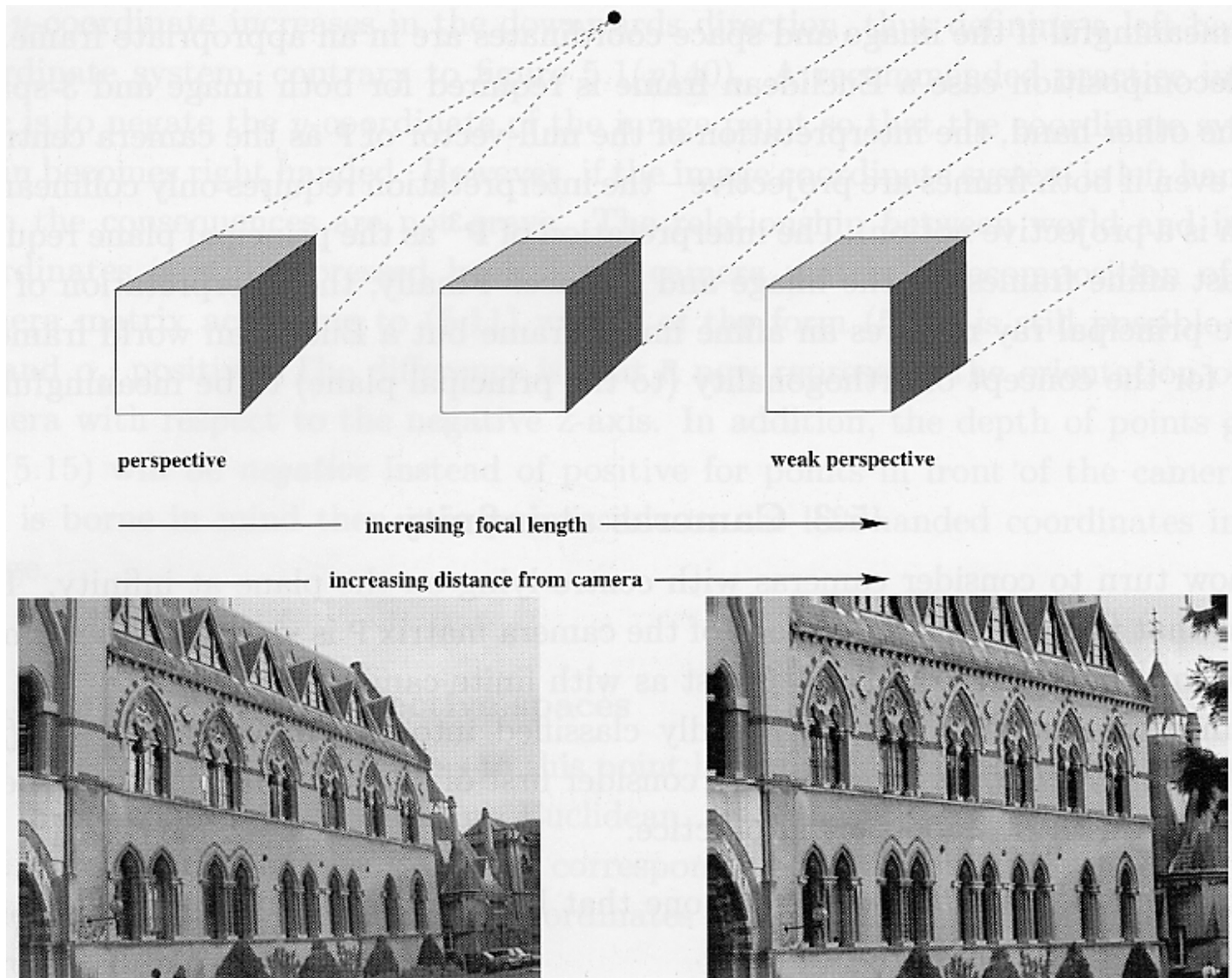


standard



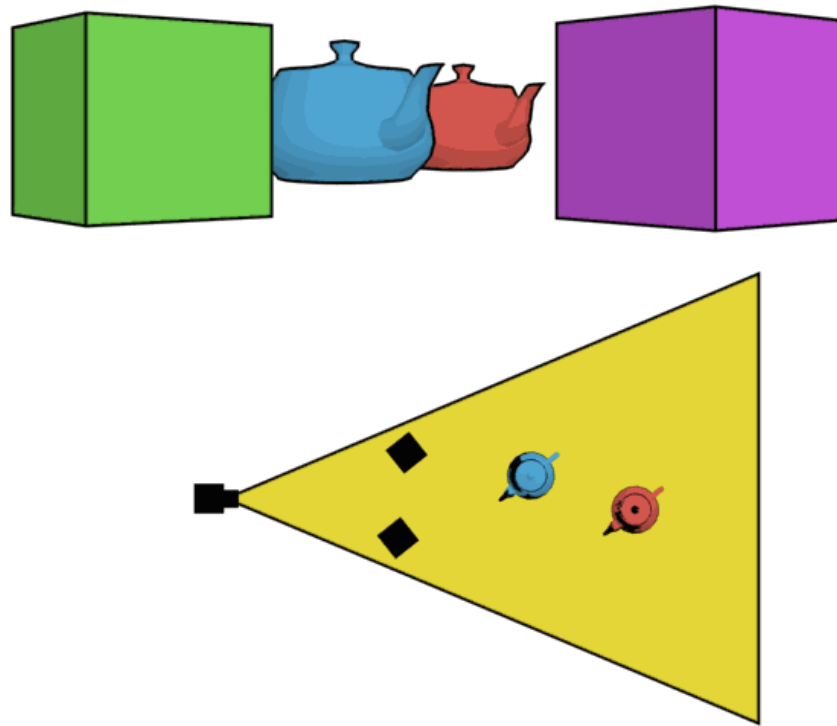
telephoto

Approximating an orthographic camera



The dolly zoom

- Continuously adjusting the focal length while the camera moves away from (or towards) the subject



http://en.wikipedia.org/wiki/Dolly_zoom

The dolly zoom

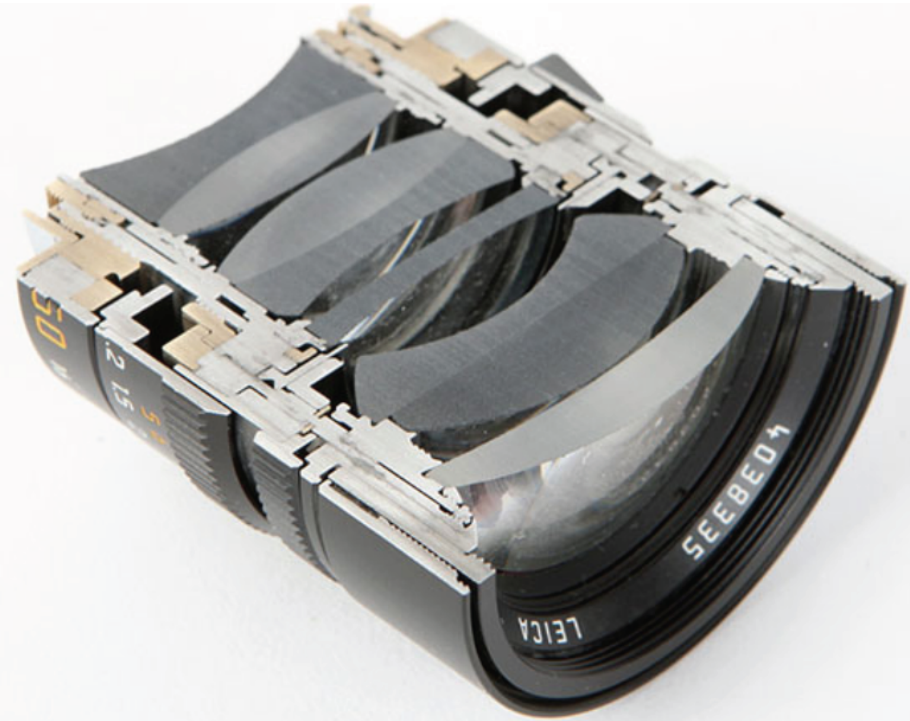
- Continuously adjusting the focal length while the camera moves away from (or towards) the subject
- “The Vertigo shot”



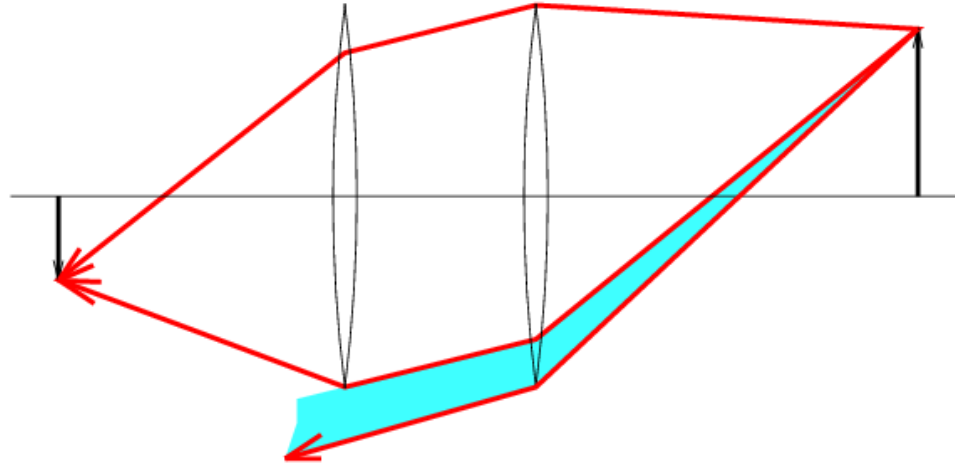
[Example of dolly zoom from *Goodfellas*](#) (YouTube)

[Example of dolly zoom from *La Haine*](#) (YouTube)

Real lenses

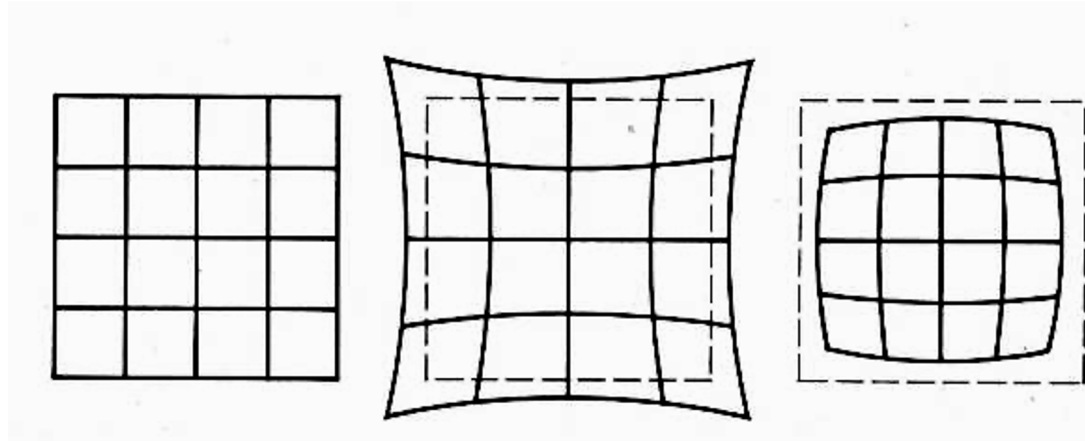


Lens flaws: Vignetting



Radial Distortion

- Caused by imperfect lenses
- Deviations are most noticeable near the edge of the lens



No distortion

Pin cushion

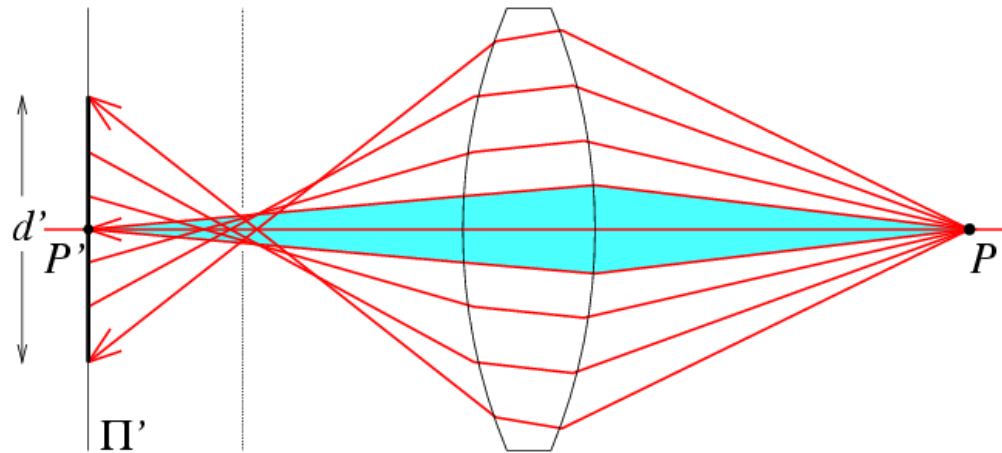
Barrel



Lens flaws: Spherical aberration

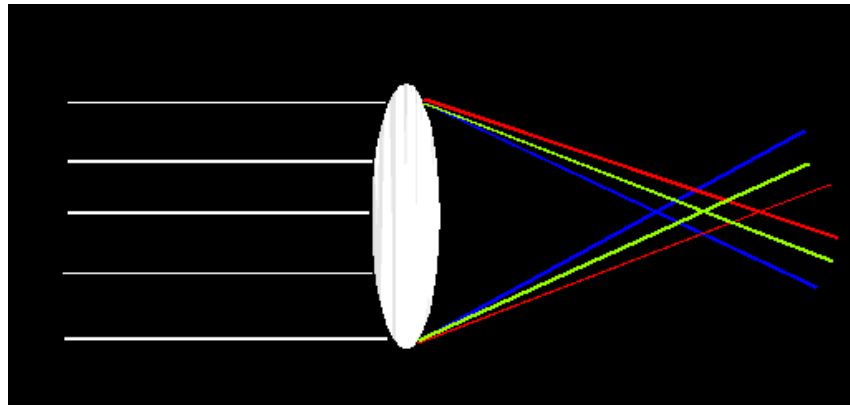
Spherical lenses don't focus light perfectly

Rays farther from the optical axis focus closer



Lens Flaws: Chromatic Aberration

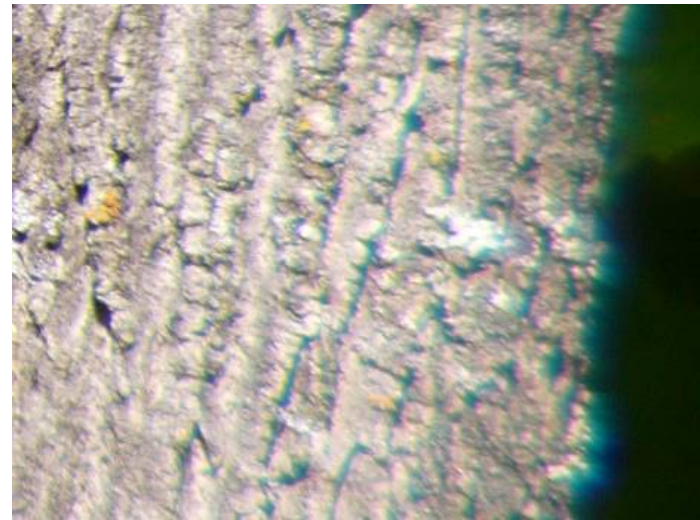
Lens has different refractive indices for different wavelengths: causes color fringing



Near Lens Center

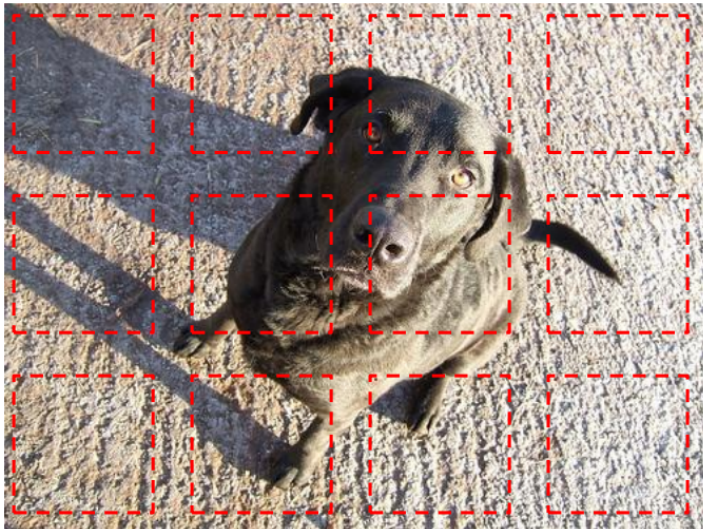


Near Lens Outer Edge



Lens Flaws: Chromatic Aberration

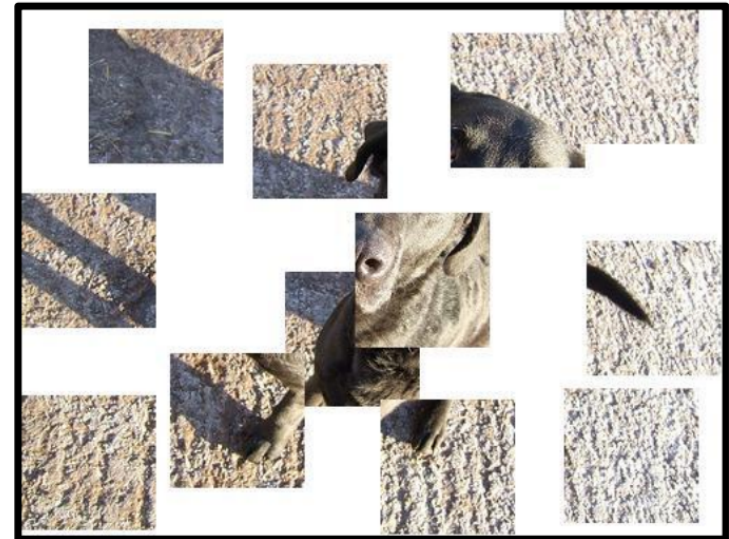
Researchers tried teaching a network about objects by forcing it to assemble jigsaws.



Initial layout, with sampled patches in red



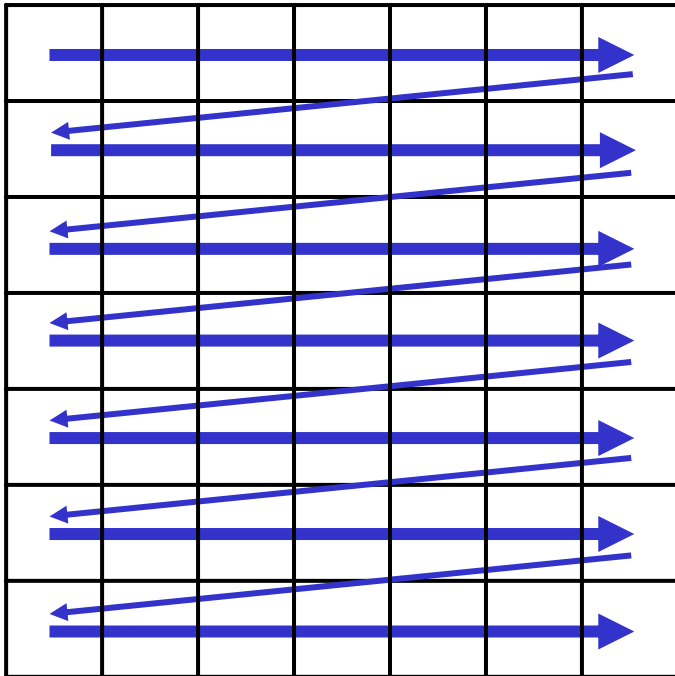
Image layout
is discarded



We can recover image layout automatically

Rolling Shutter

Rolling Shutter: pixels read in sequence
Can get global reading, but \$\$\$

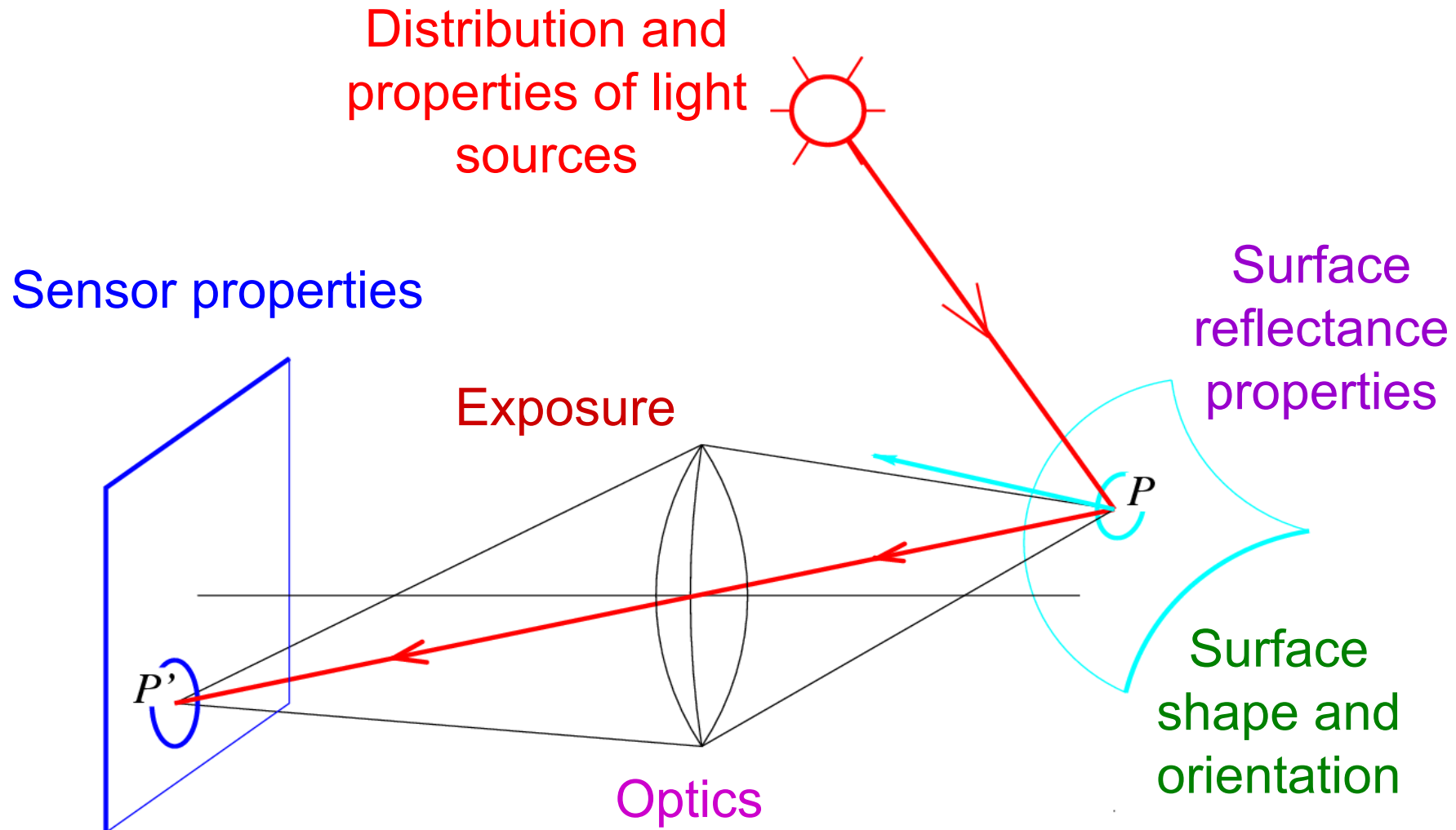


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Image formation

What determines the brightness of an image pixel?



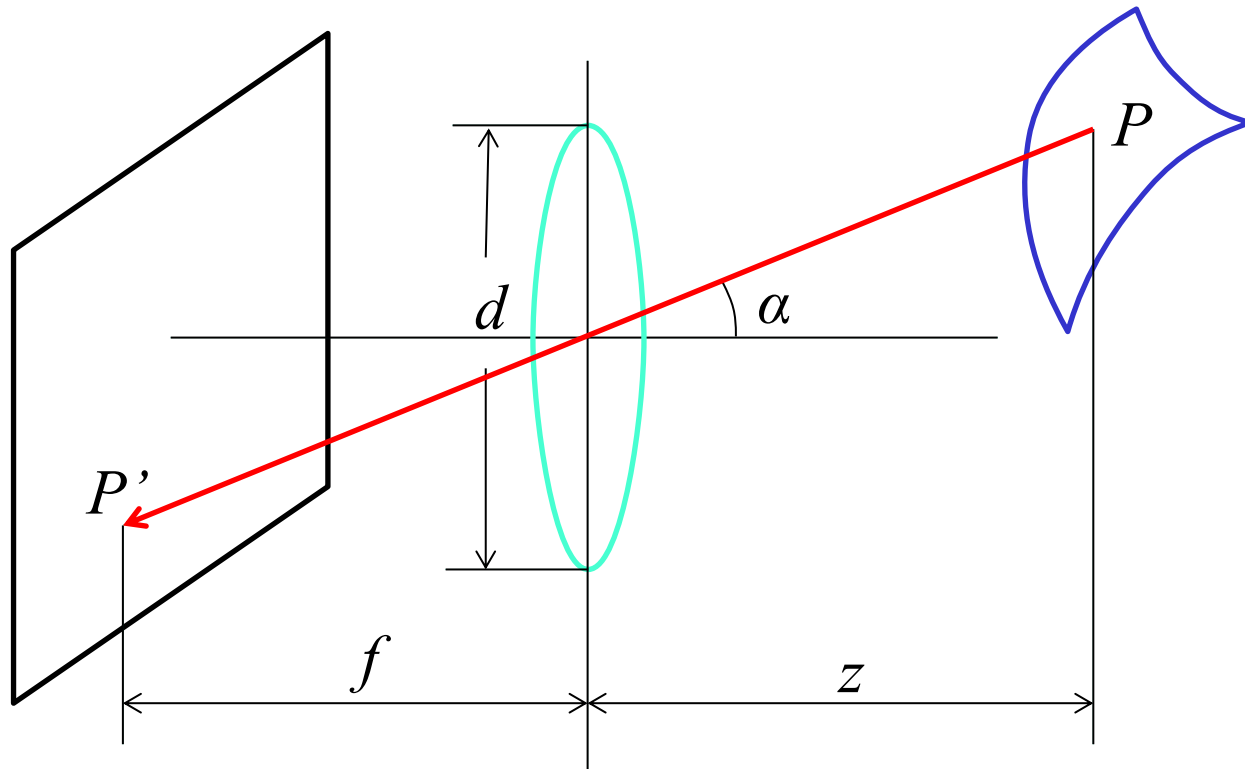
Fundamental radiometric relation

L: Radiance emitted from P toward P'

- Energy carried by a ray (Watts per sq. meter per steradian)

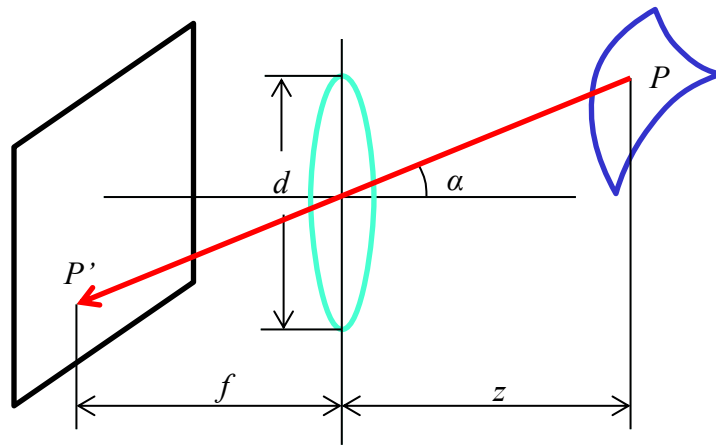
E: Irradiance falling on P' from the lens

- Energy arriving at a surface (Watts per sq. meter)



What is the relationship between *E* and *L*?

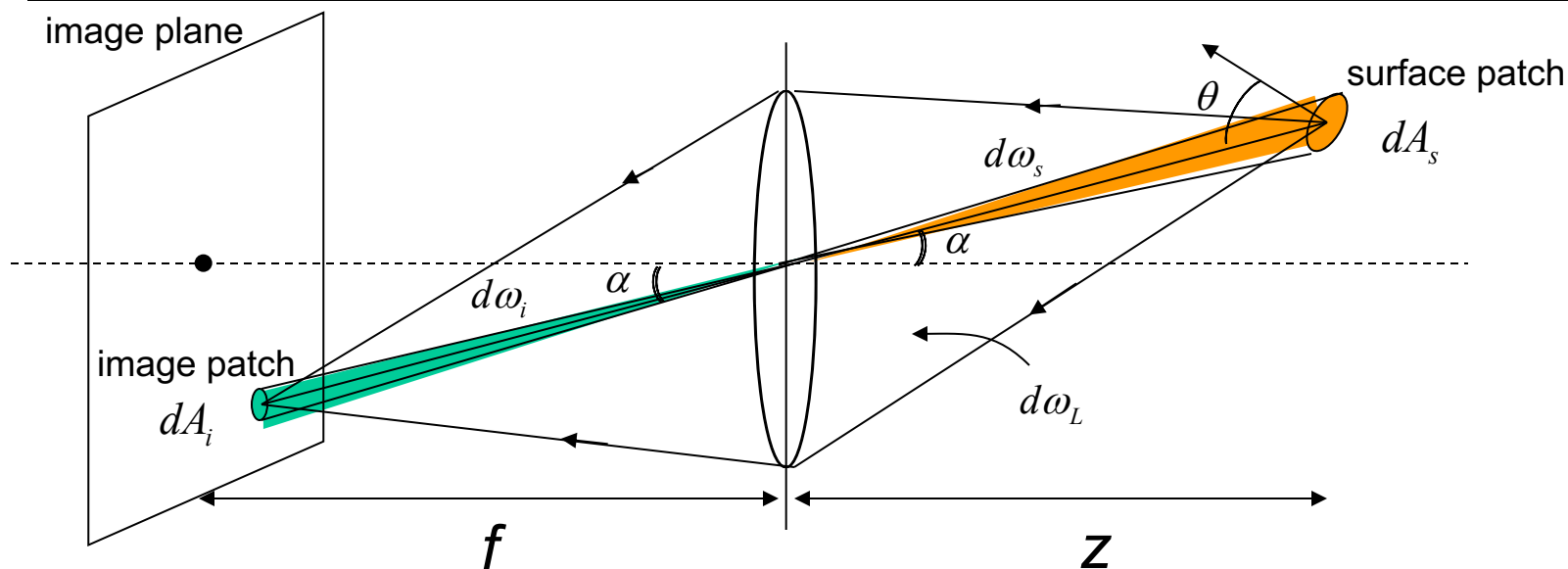
Fundamental radiometric relation



$$E = \left[\frac{\pi}{4} \left(\frac{d}{f} \right)^2 \cos^4 \alpha \right] L$$

- Image irradiance is linearly related to scene radiance
- Irradiance is proportional to the area of the lens and inversely proportional to the squared distance between the lens and the image plane
- The irradiance falls off as the angle between the viewing ray and the optical axis increases

Relation between Image Irradiance E and Scene Radiance L



- Solid angles of the double cone (orange and green):

$$d\omega_i = d\omega_s \quad \frac{dA_i \cos \alpha}{(f / \cos \alpha)^2} = \frac{dA_s \cos \theta}{(z / \cos \alpha)^2}$$

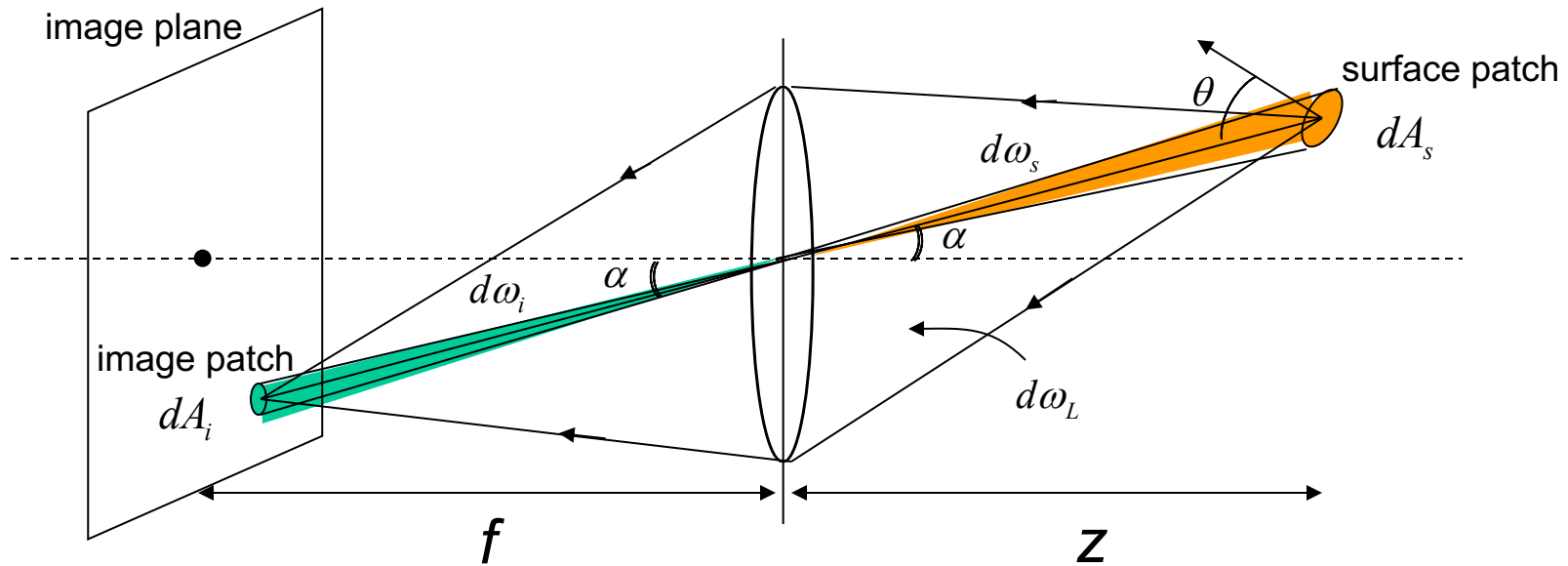
$$\frac{dA_s}{dA_i} = \frac{\cos \alpha}{\cos \theta} \left(\frac{z}{f} \right)^2$$

- Solid angle subtended by lens:

$$d\omega_L = \frac{\pi d^2}{4} \frac{\cos \alpha}{(z / \cos \alpha)^2} \rightarrow (2)$$

(1)

Relation between Image Irradiance E and Scene Radiance L



- Flux received by lens from dA_s = Flux projected onto image dA_i

$$L (dA_s \cos \theta) d\omega_L = E dA_i \quad \rightarrow (3)$$

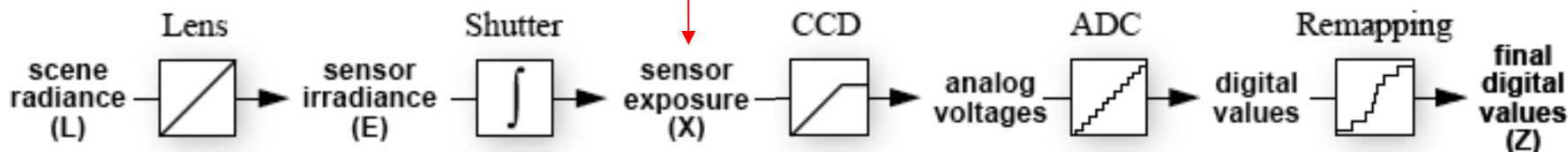
- From (1), (2), and (3):
$$E = L \frac{\pi}{4} \left(\frac{d}{f} \right)^2 \cos \alpha^4$$

• Image irradiance is proportional to Scene Radiance!

• Small field of view \rightarrow Effects of 4th power of cosine are small.

From light rays to pixel values

$$X = E \cdot \Delta t$$



$$E = \left[\frac{\pi}{4} \left(\frac{d}{f} \right)^2 \cos^4 \alpha \right] L$$

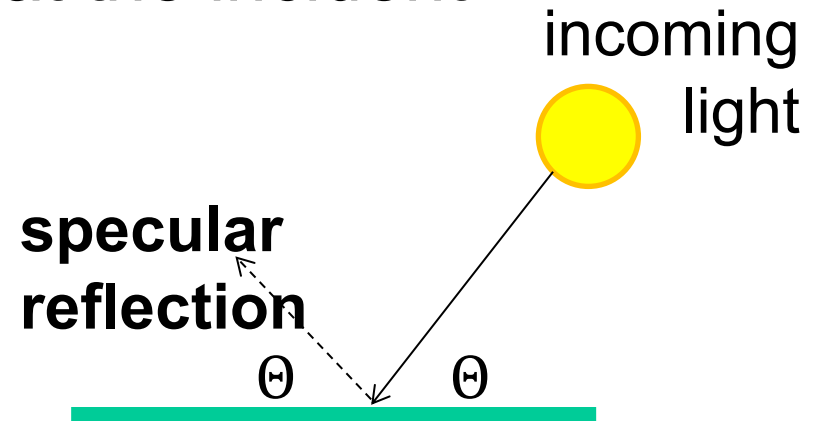
$$Z = f(E \cdot \Delta t)$$

- Camera response function: the mapping f from irradiance to pixel values
 - Useful if we want to estimate material properties
 - Enables us to create *high dynamic range (HDR) images*
 - Classic reference: P. E. Debevec and J. Malik, [Recovering High Dynamic Range Radiance Maps from Photographs](#), SIGGRAPH 97

Basic models of reflection

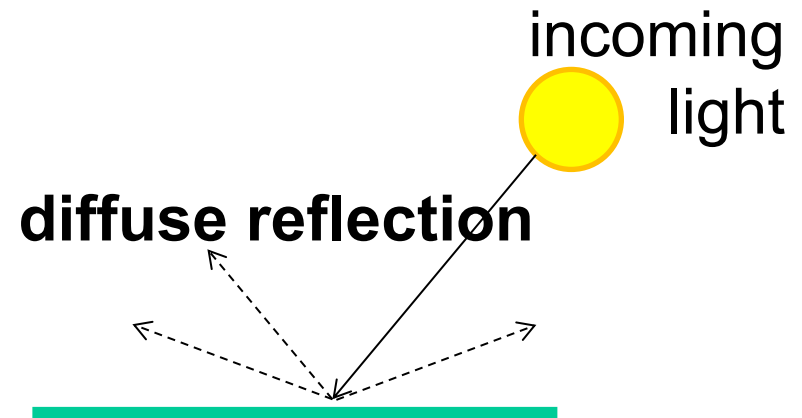
Specular: light bounces off at the incident angle

- E.g., mirror



Diffuse: light scatters in all directions

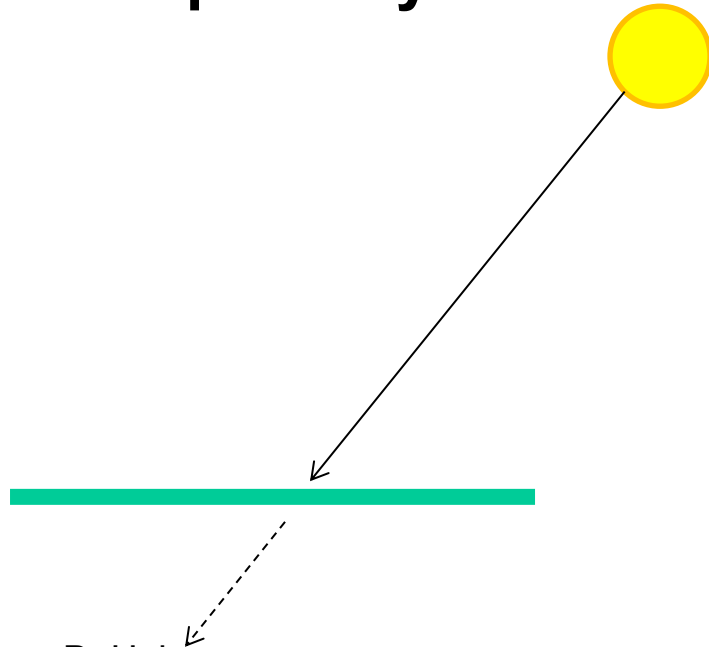
- E.g., brick, cloth, rough wood



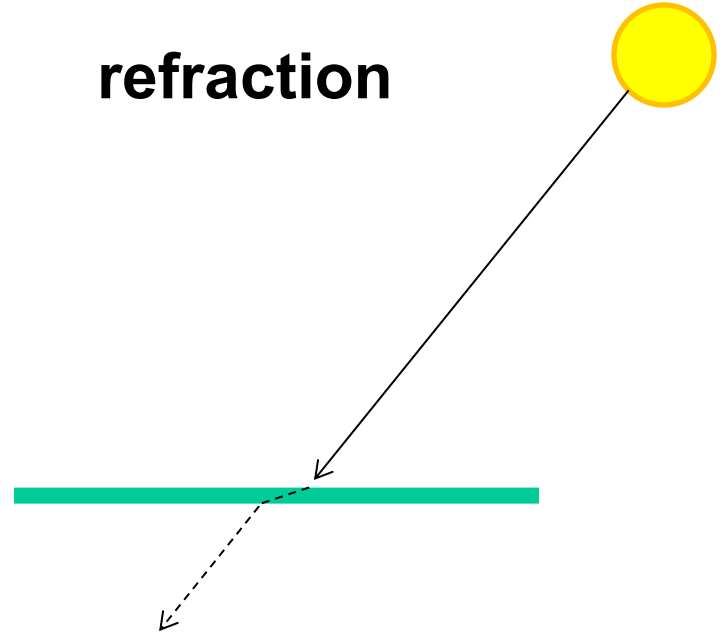
Other possible effects



transparency light source



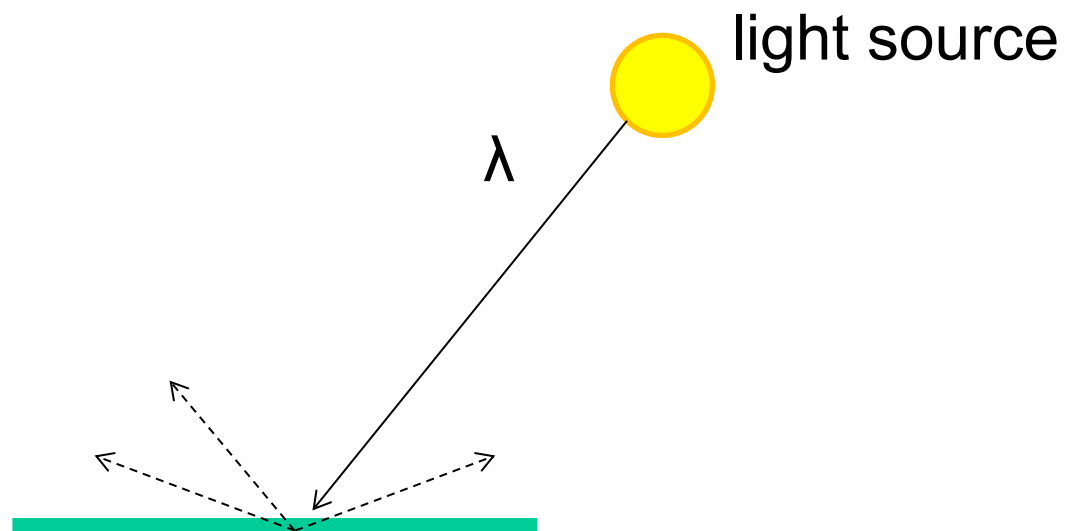
refraction light source



Other possible effects



**subsurface
scattering**

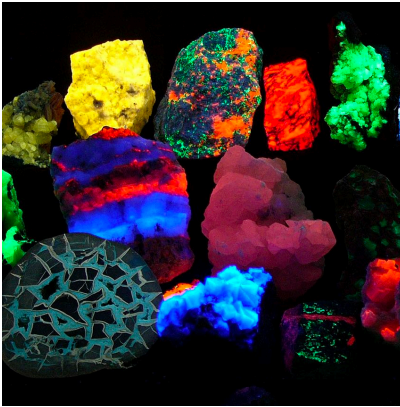


Slide from D. Hoiem

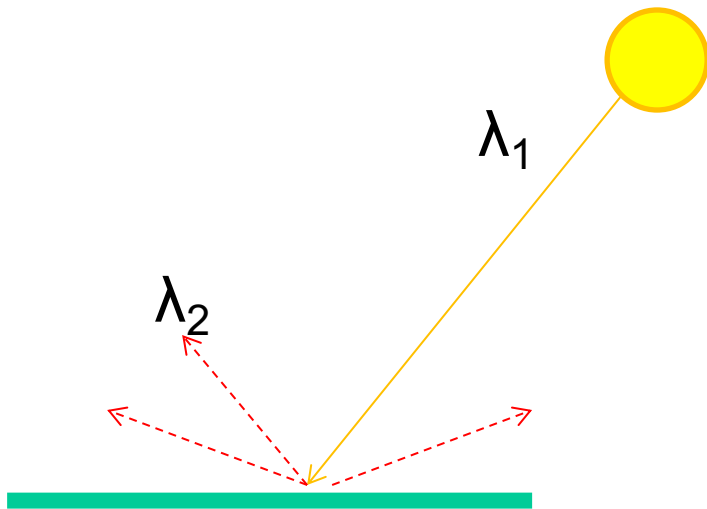
https://en.wikipedia.org/wiki/Subsurface_scattering#/media/File:Skin_Subsurface_Scattering.jpg

Other possible effects

fluorescence



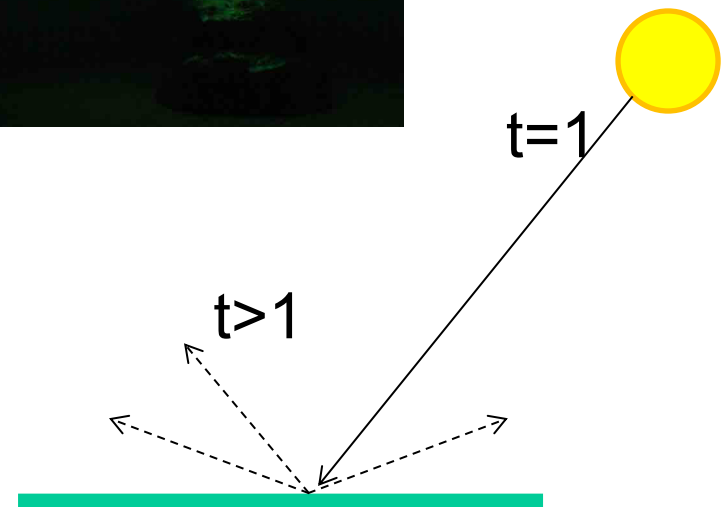
light source



phosphorescence



light source

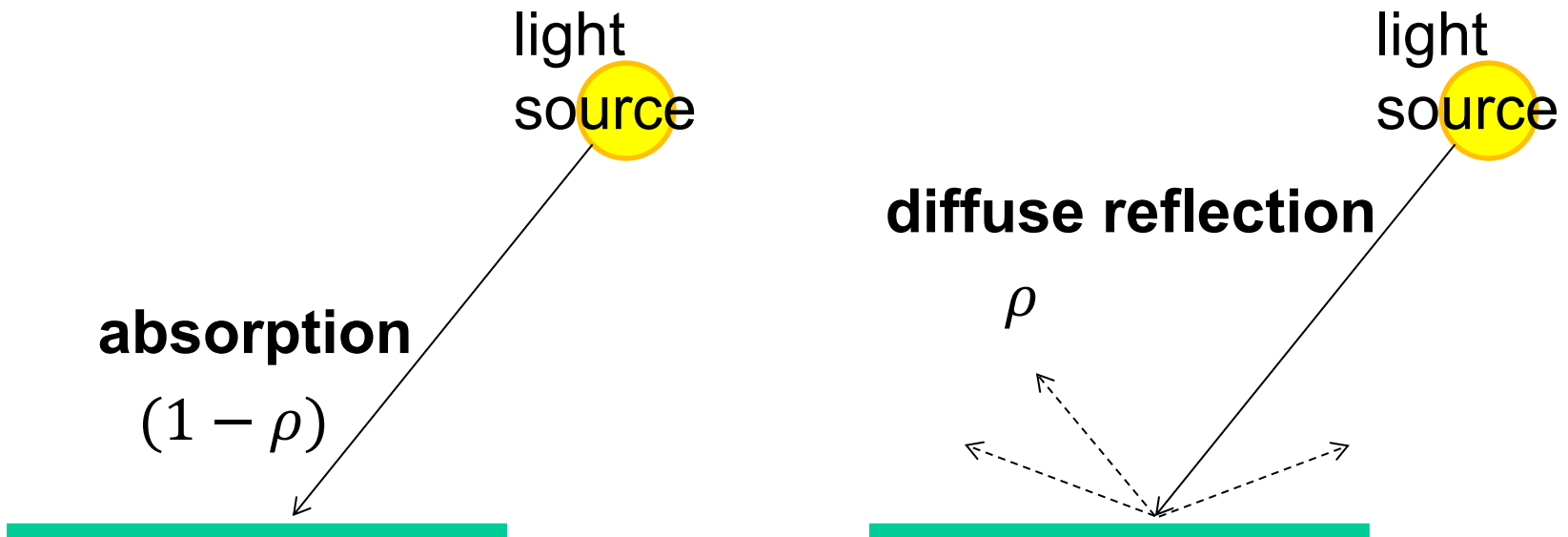


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Lambertian reflectance model

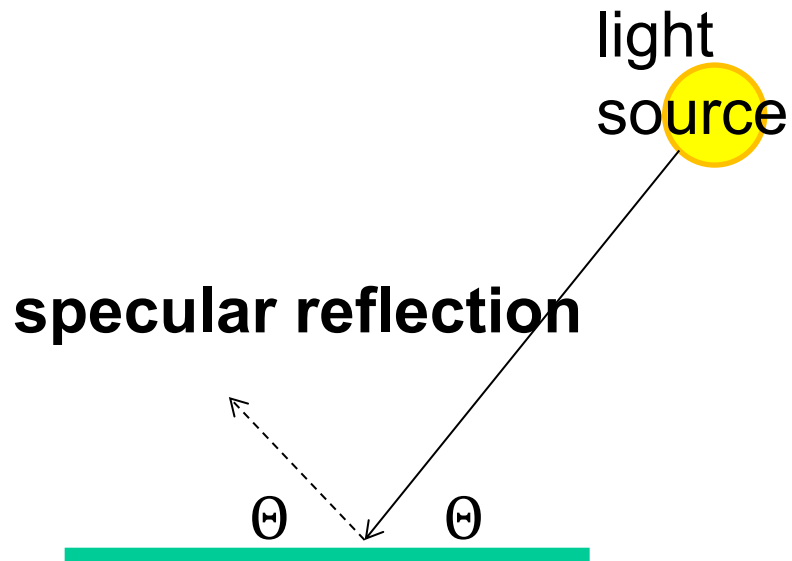
Some light is absorbed (function of albedo ρ)
Remaining light is scattered (diffuse reflection)
Examples: soft cloth, concrete, matte paints



Specular Reflection

Reflected direction depends on light orientation and surface normal

- E.g., mirrors are fully specular
- Most surfaces can be modeled with a mixture of diffuse and specular components



Flickr, by suzysputnik



Flickr, by piratejohnny

Most surfaces have both specular and ~~diffuse components~~

Specularity = spot where specular reflection dominates (typically reflects light source)



Photo: northcountryhardwoodfloors.com



Typically, specular component is small

Intensity and Surface Orientation

Intensity depends on illumination angle because less light comes in at oblique angles.

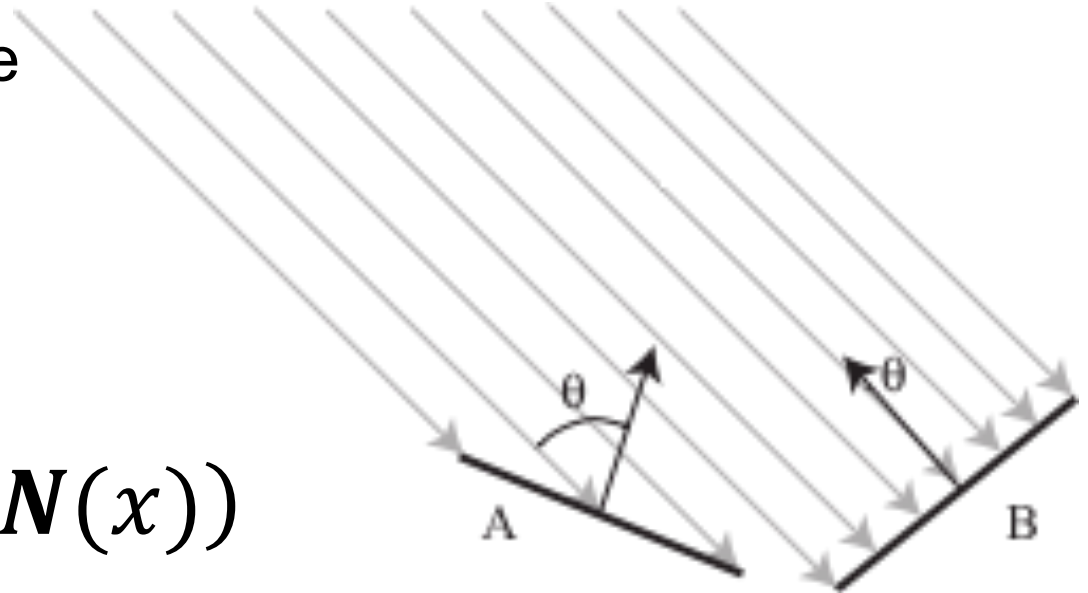
ρ = albedo

\mathbf{S} = directional source

\mathbf{N} = surface normal

I = reflected intensity

$$I(x) = \rho(x)(\mathbf{S} \cdot \mathbf{N}(x))$$



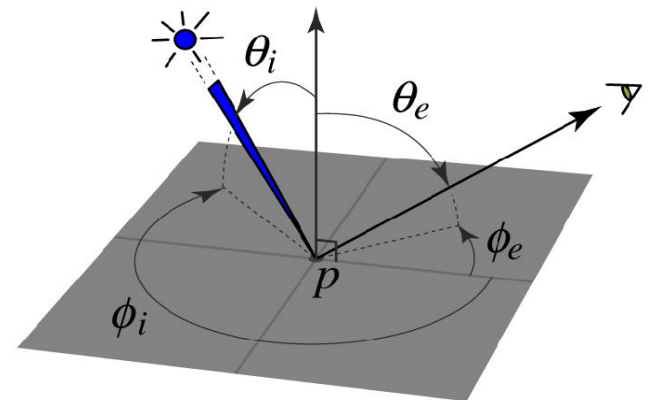
The interaction of light and surfaces

What happens when a light ray hits a point on an object?

- Some of the light gets **absorbed**
 - converted to other forms of energy (e.g., heat)
- Some gets **transmitted** through the object
 - possibly bent, through refraction
 - or scattered inside the object (subsurface scattering)
- Some gets **reflected**
 - possibly in multiple directions at once
- Really complicated things can happen
 - fluorescence

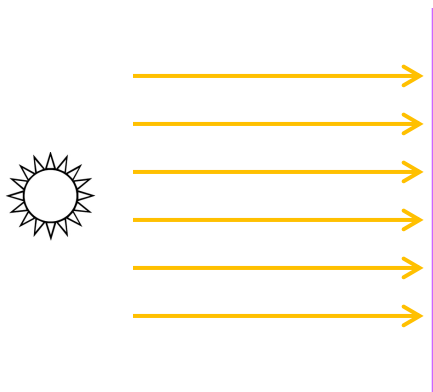
Bidirectional reflectance distribution function (BRDF)

- How bright a surface appears when viewed from one direction when light falls on it from another
- Definition: ratio of the radiance in the emitted direction to irradiance in the incident direction

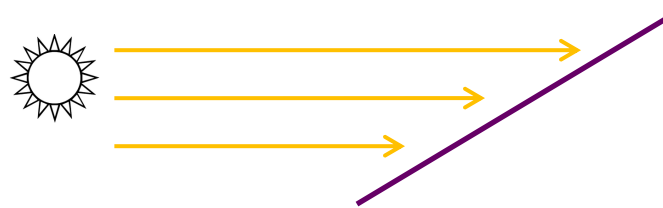


Diffuse reflection

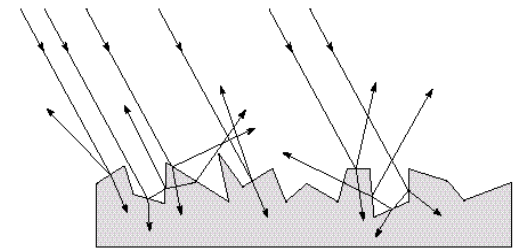
- Light is reflected equally in all directions
 - Dull, matte surfaces like chalk or latex paint
 - Microfacets scatter incoming light randomly
- Brightness of the surface depends on the incidence of illumination



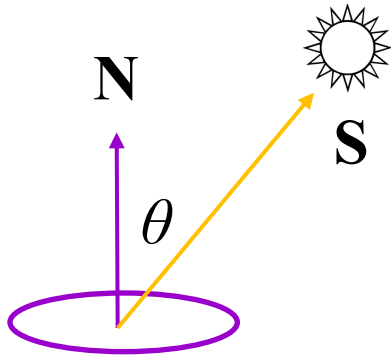
brighter



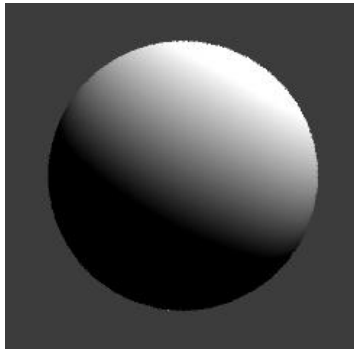
darker



Diffuse reflection: Lambert's law



$$B = \rho \mathbf{N} \cdot \mathbf{S}$$
$$= \rho \|\mathbf{S}\| \cos \theta$$



B : radiosity (total power leaving the surface per unit area)

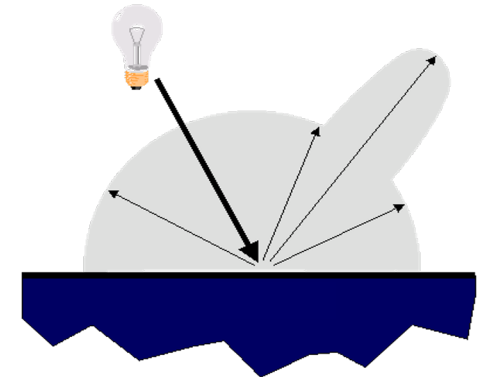
ρ : albedo (fraction of incident irradiance reflected by the surface)

N : unit normal

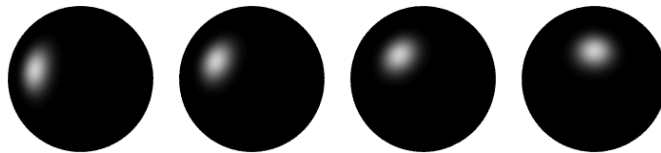
S : source vector (magnitude proportional to intensity of the source)

Specular reflection

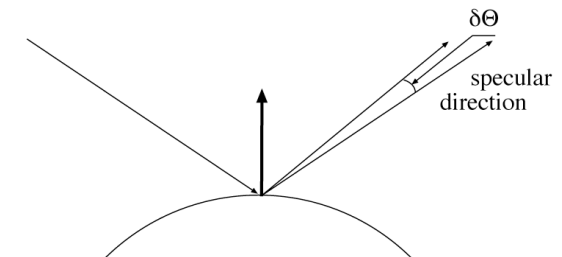
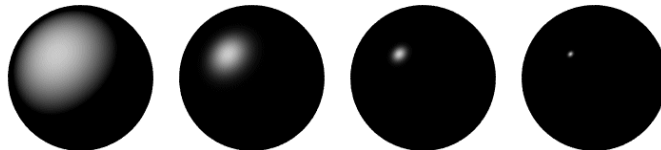
- Radiation arriving along a source direction leaves along the **specular direction** (source direction reflected about normal)
- On real surfaces, energy usually goes into a lobe of directions
- **Phong model:** reflected energy falls off with $\cos^n(\delta\theta)$



Moving the light source



Changing the exponent



Specular reflection



[Picture source](#)

Photometric stereo (shape from shading)

- Can we reconstruct the shape of an object based on shading cues?



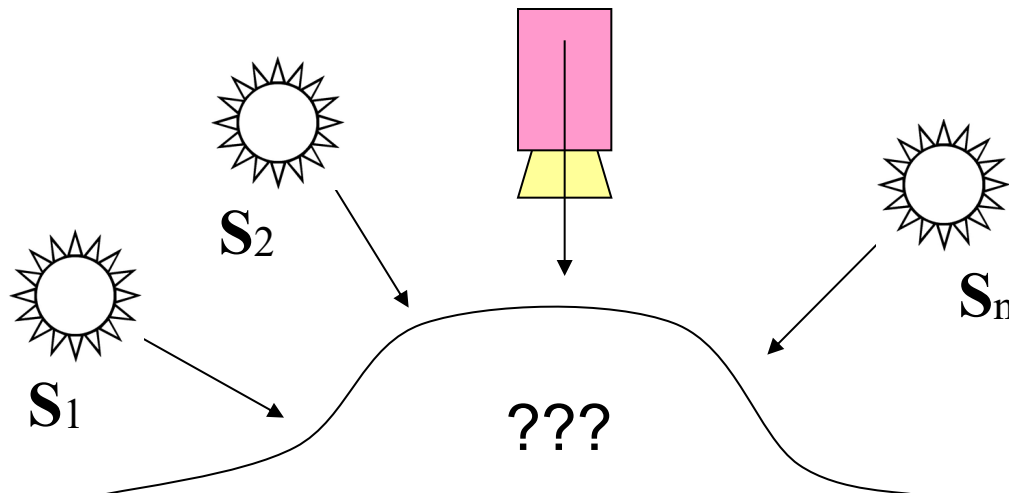
Luca della Robbia,
Cantoria, 1438

Photometric stereo

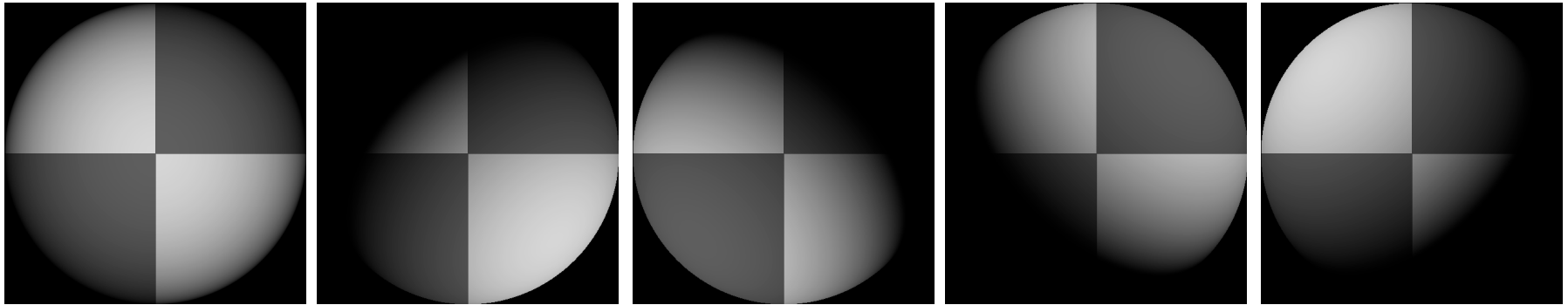
Assume:

- A Lambertian object
- A *local shading model* (each point on a surface receives light only from sources visible at that point)
- A set of *known* light source directions
- A set of pictures of an object, obtained in exactly the same camera/object configuration but using different sources
- Orthographic projection

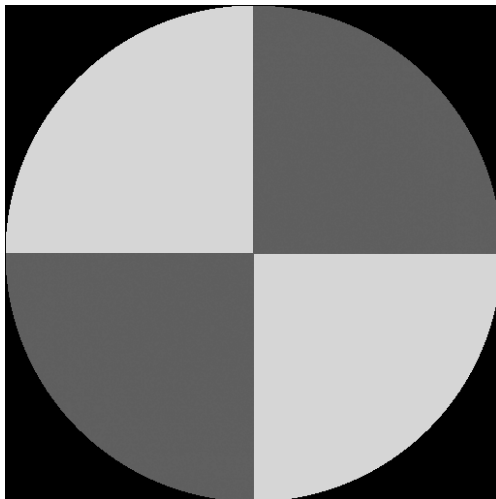
Goal: reconstruct object shape and albedo



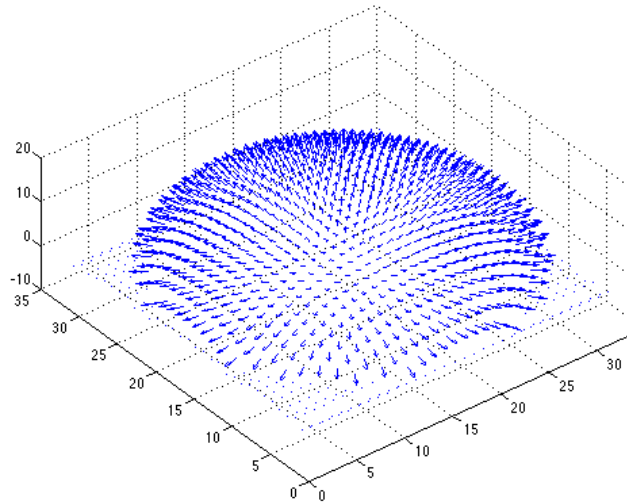
Example 1



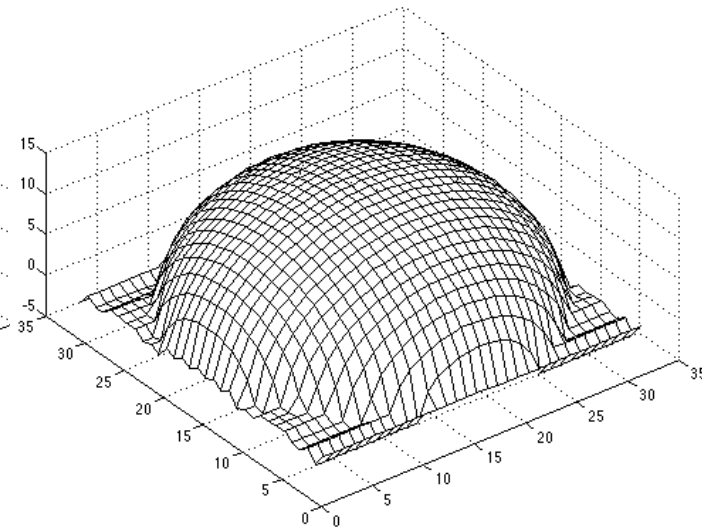
Recovered
albedo



Recovered normal
field



Recovered surface
model



Example 2

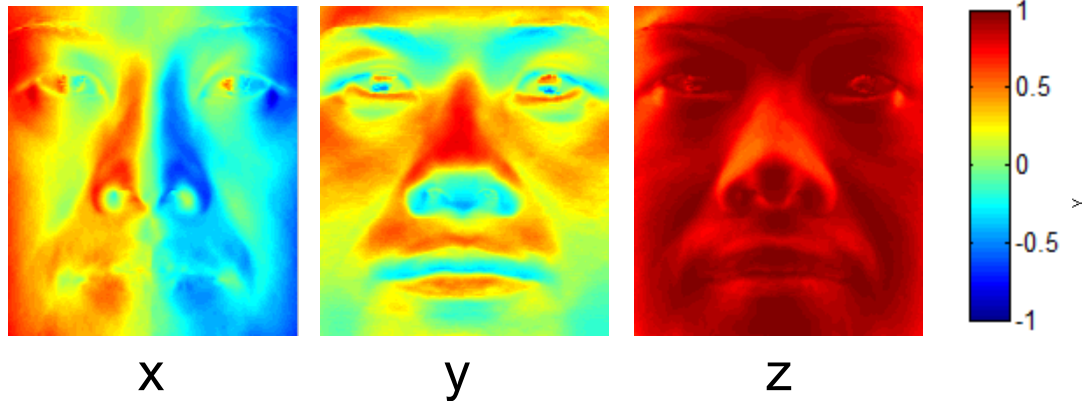
Input



Recovered albedo



Recovered normal field



Recovered surface model

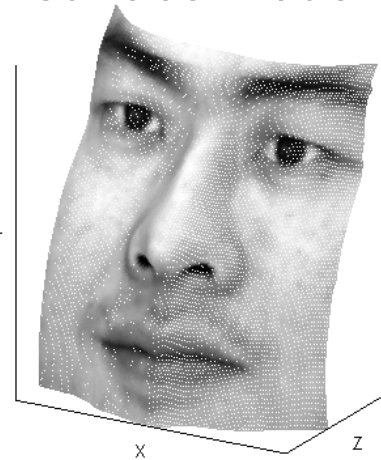


Image model

- **Known:** source vectors S_j and pixel values $I_j(x,y)$
- **Unknown:** surface normal $\mathbf{N}(x,y)$ and albedo $\rho(x,y)$

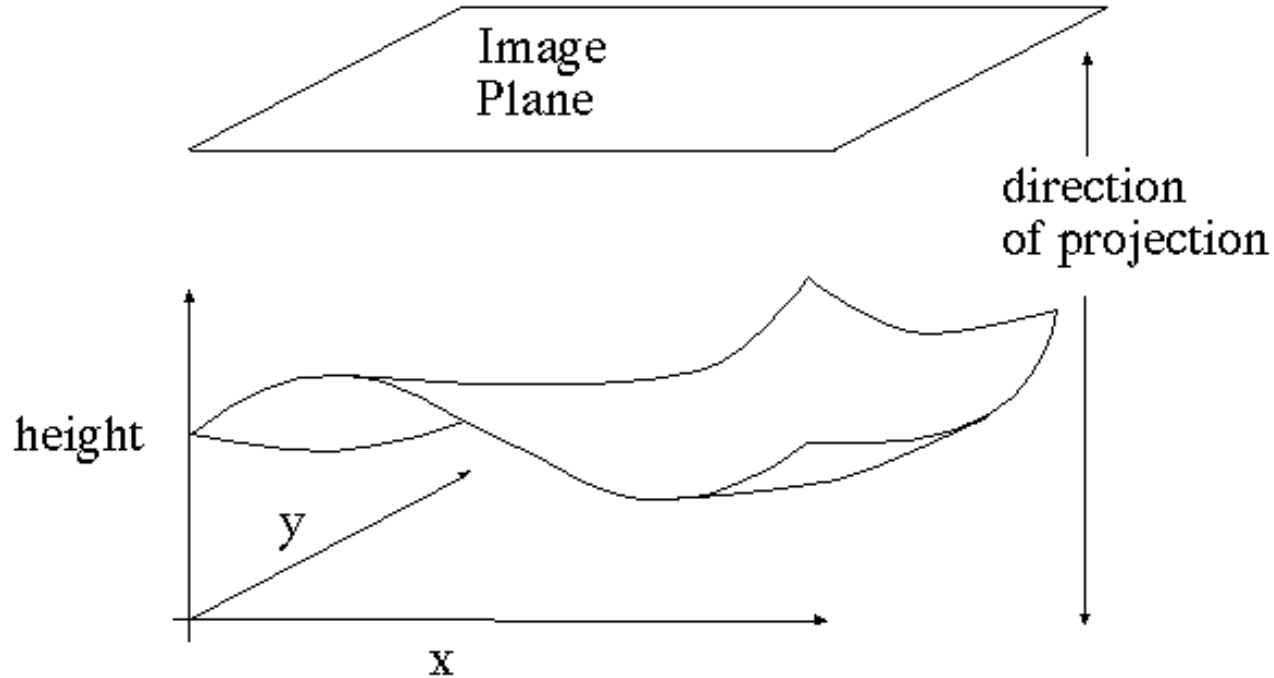


Image model

- **Known:** source vectors \mathbf{S}_j and pixel values $I_j(x,y)$
- **Unknown:** surface normal $\mathbf{N}(x,y)$ and albedo $\rho(x,y)$
- Assume that the response function of the camera is a linear scaling by a factor of k
- Lambert's law:

$$\begin{aligned} I_j(x, y) &= k \rho(x, y) (\mathbf{N}(x, y) \cdot \mathbf{S}_j) \\ &= (\rho(x, y) \mathbf{N}(x, y)) \cdot (k \mathbf{S}_j) \\ &= \mathbf{g}(x, y) \cdot \mathbf{V}_j \end{aligned}$$

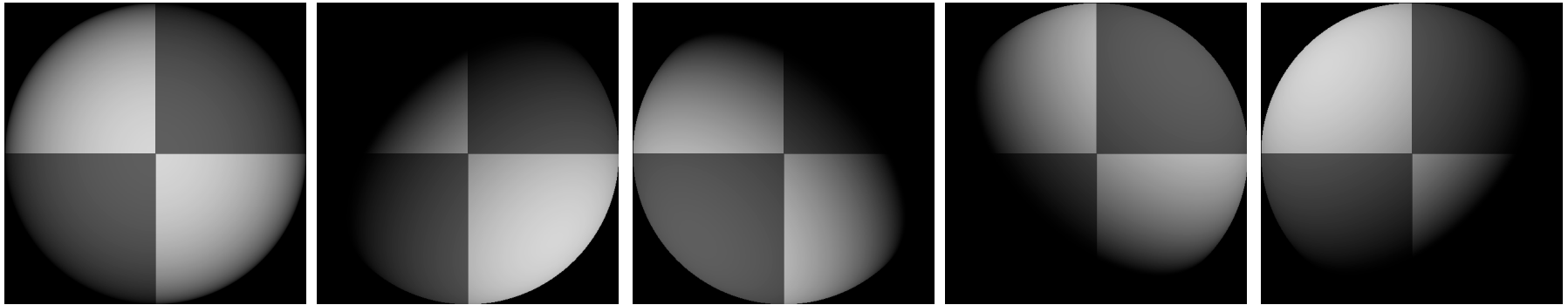
Least squares problem

- For each pixel, set up a linear system:

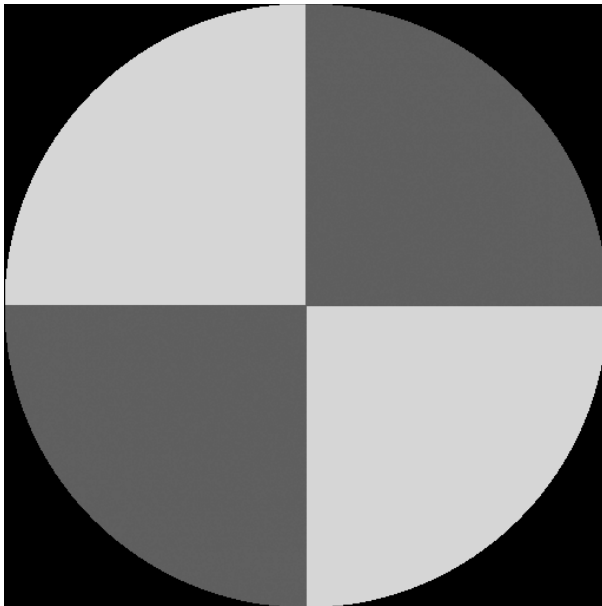
$$\begin{array}{c} \left[\begin{array}{c} I_1(x, y) \\ I_2(x, y) \\ \vdots \\ I_n(x, y) \end{array} \right] = \left[\begin{array}{c} \mathbf{V}_1^T \\ \mathbf{V}_2^T \\ \vdots \\ \mathbf{V}_n^T \end{array} \right] \mathbf{g}(x, y) \\ \begin{array}{c} | \\ (n \times 1) \\ \text{known} \end{array} \quad \begin{array}{c} | \\ (n \times 3) \\ \text{known} \end{array} \quad \begin{array}{c} | \\ (3 \times 1) \\ \text{unknown} \end{array} \end{array}$$

- Obtain least-squares solution for $\mathbf{g}(x, y)$
(which we defined as $\mathbf{N}(x, y) \rho(x, y)$)
- Since $\mathbf{N}(x, y)$ is the unit normal, $\rho(x, y)$ is given by the magnitude of $\mathbf{g}(x, y)$
- Finally, $\mathbf{N}(x, y) = \mathbf{g}(x, y) / \rho(x, y)$

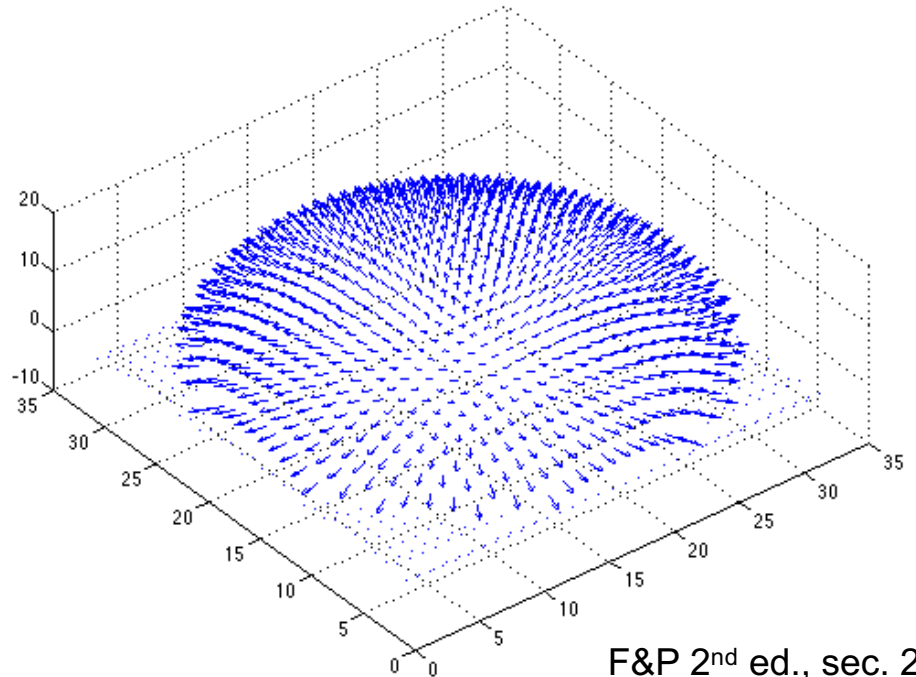
Synthetic example



Recovered albedo



Recovered normal field



Recovering a surface from normals

Recall the surface is written as

$$(x, y, f(x, y))$$

This means the normal has the form:

$$\mathbf{N}(x, y) = \frac{1}{\sqrt{f_x^2 + f_y^2 + 1}} \begin{pmatrix} f_x \\ f_y \\ 1 \end{pmatrix}$$

If we write the estimated vector g as

$$\mathbf{g}(x, y) = \begin{pmatrix} g_1(x, y) \\ g_2(x, y) \\ g_3(x, y) \end{pmatrix}$$

Then we obtain values for the partial derivatives of the surface:

$$f_x(x, y) = g_1(x, y) / g_3(x, y)$$

$$f_y(x, y) = g_2(x, y) / g_3(x, y)$$

Recovering a surface from normals

We can now recover the surface height at any point by integration along some path, e.g.

$$f(x, y) = \int_0^x f_x(s, 0) ds + \int_0^y f_y(x, t) dt + C$$

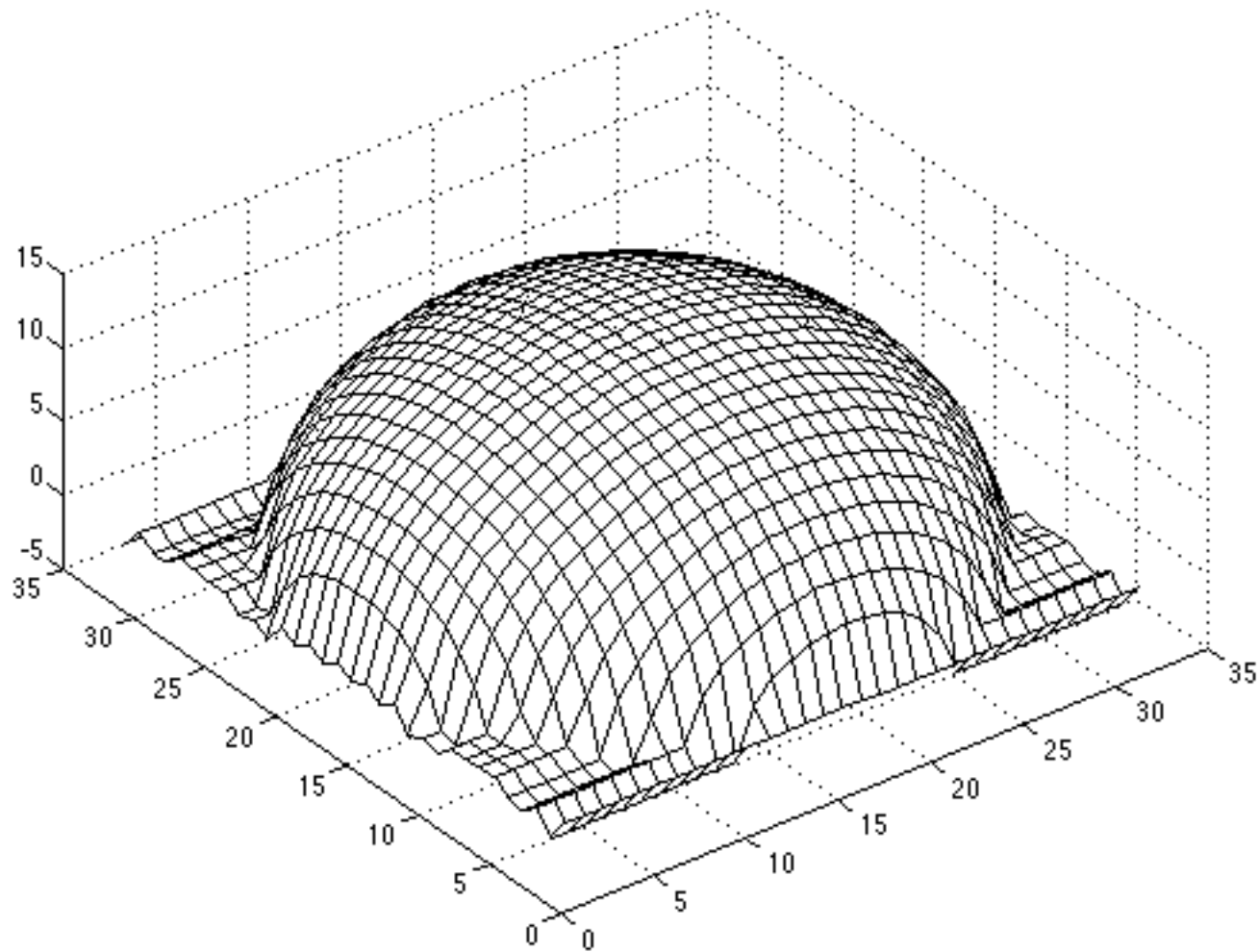
(for robustness, should take integrals over many different paths and average the results)

Integrability: for the surface f to exist, the mixed second partial derivatives must be equal:

$$\frac{\partial}{\partial y} (g_1(x, y) / g_3(x, y)) = \frac{\partial}{\partial x} (g_2(x, y) / g_3(x, y))$$

(in practice, they should at least be similar)

Surface recovered by integration



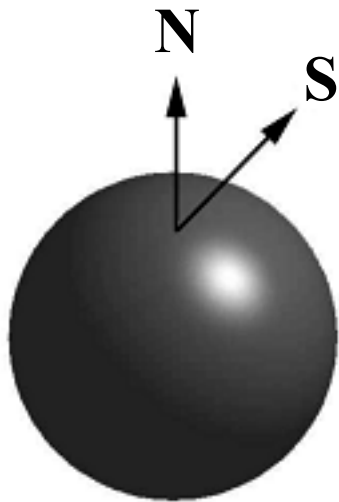
Limitations

- Orthographic camera model
- Simplistic reflectance and lighting model
- No shadows
- No interreflections
- No missing data
- Integration is tricky

Finding the direction of the light source

$$I(x,y) = \mathbf{N}(x,y) \cdot \mathbf{S}(x,y)$$

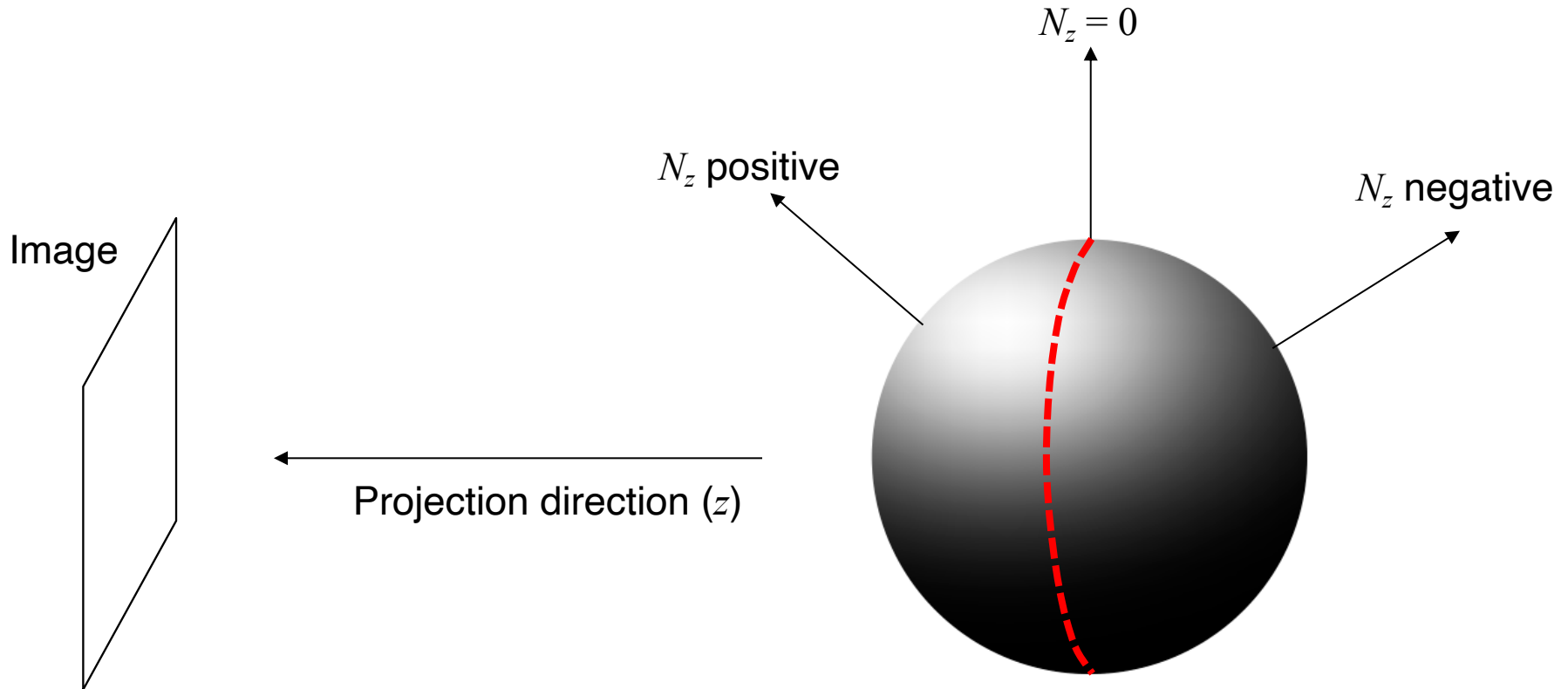
Full 3D case:



$$\begin{pmatrix} N_x(x_1, y_1) & N_y(x_1, y_1) & N_z(x_1, y_1) \\ N_x(x_2, y_2) & N_y(x_2, y_2) & N_z(x_2, y_2) \\ \vdots & \vdots & \vdots \\ N_x(x_n, y_n) & N_y(x_n, y_n) & N_z(x_n, y_n) \end{pmatrix} \begin{pmatrix} S_x \\ S_y \\ S_z \end{pmatrix} = \begin{pmatrix} I(x_1, y_1) \\ I(x_2, y_2) \\ \vdots \\ I(x_n, y_n) \end{pmatrix}$$

Finding the direction of the light source

Consider points on the *occluding contour*:

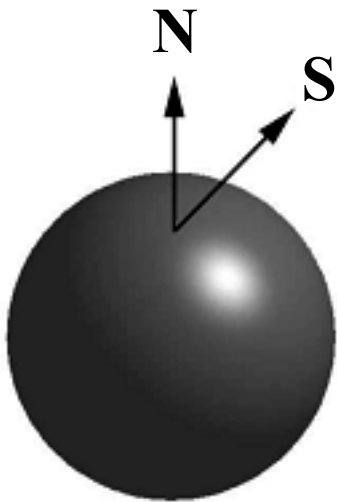


Finding the direction of the light source

$$I(x,y) = \mathbf{N}(x,y) \cdot \mathbf{S}(x,y)$$

Full 3D case:

$$\begin{pmatrix} N_x(x_1, y_1) & N_y(x_1, y_1) & N_z(x_1, y_1) \\ N_x(x_2, y_2) & N_y(x_2, y_2) & N_z(x_2, y_2) \\ \vdots & \vdots & \vdots \\ N_x(x_n, y_n) & N_y(x_n, y_n) & N_z(x_n, y_n) \end{pmatrix} \begin{pmatrix} S_x \\ S_y \\ S_z \end{pmatrix} = \begin{pmatrix} I(x_1, y_1) \\ I(x_2, y_2) \\ \vdots \\ I(x_n, y_n) \end{pmatrix}$$



For points on the *occluding contour*, $N_z = 0$:

$$\begin{pmatrix} N_x(x_1, y_1) & N_y(x_1, y_1) \\ N_x(x_2, y_2) & N_y(x_2, y_2) \\ \vdots & \vdots \\ N_x(x_n, y_n) & N_y(x_n, y_n) \end{pmatrix} \begin{pmatrix} S_x \\ S_y \end{pmatrix} = \begin{pmatrix} I(x_1, y_1) \\ I(x_2, y_2) \\ \vdots \\ I(x_n, y_n) \end{pmatrix}$$

Finding the direction of the light source



P. Nillius and J.-O. Eklundh, "Automatic estimation of the projected light source direction," CVPR 2001

Application: Detecting composite photos

Fake photo



Real photo



M. K. Johnson and H. Farid, [Exposing Digital Forgeries by Detecting Inconsistencies in Lighting](#), ACM Multimedia and Security Workshop, 2005.